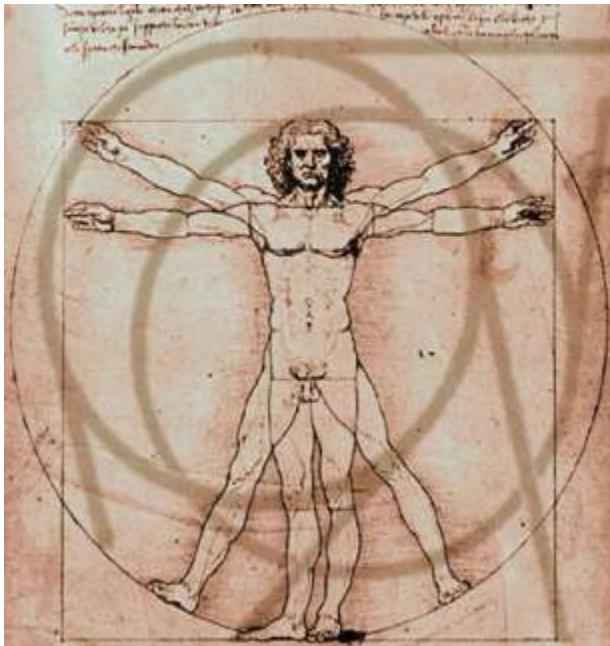
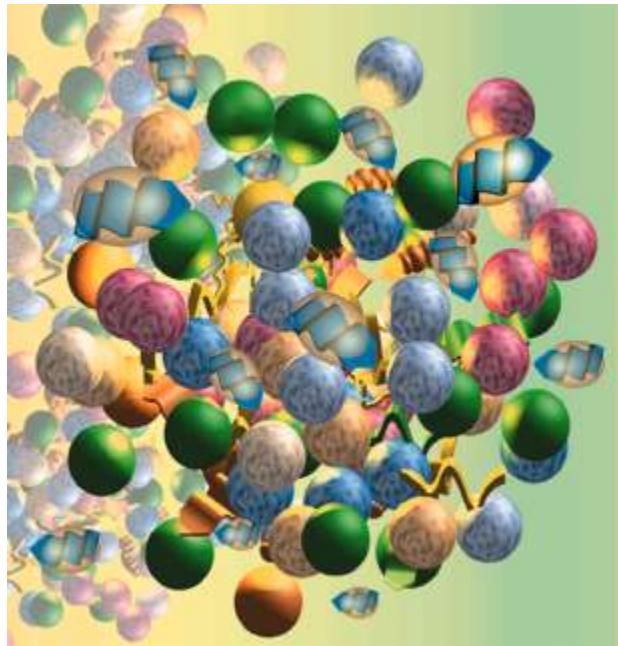


Što [još] ne znamo?



1.dio

Osnovna škola Plokite
Split, 23. travnja 2013.

Ivica Puljak
FESB – Split

Znamo li uopće nešto?

“Nitko ništa ne zna
Krhko je znanje “

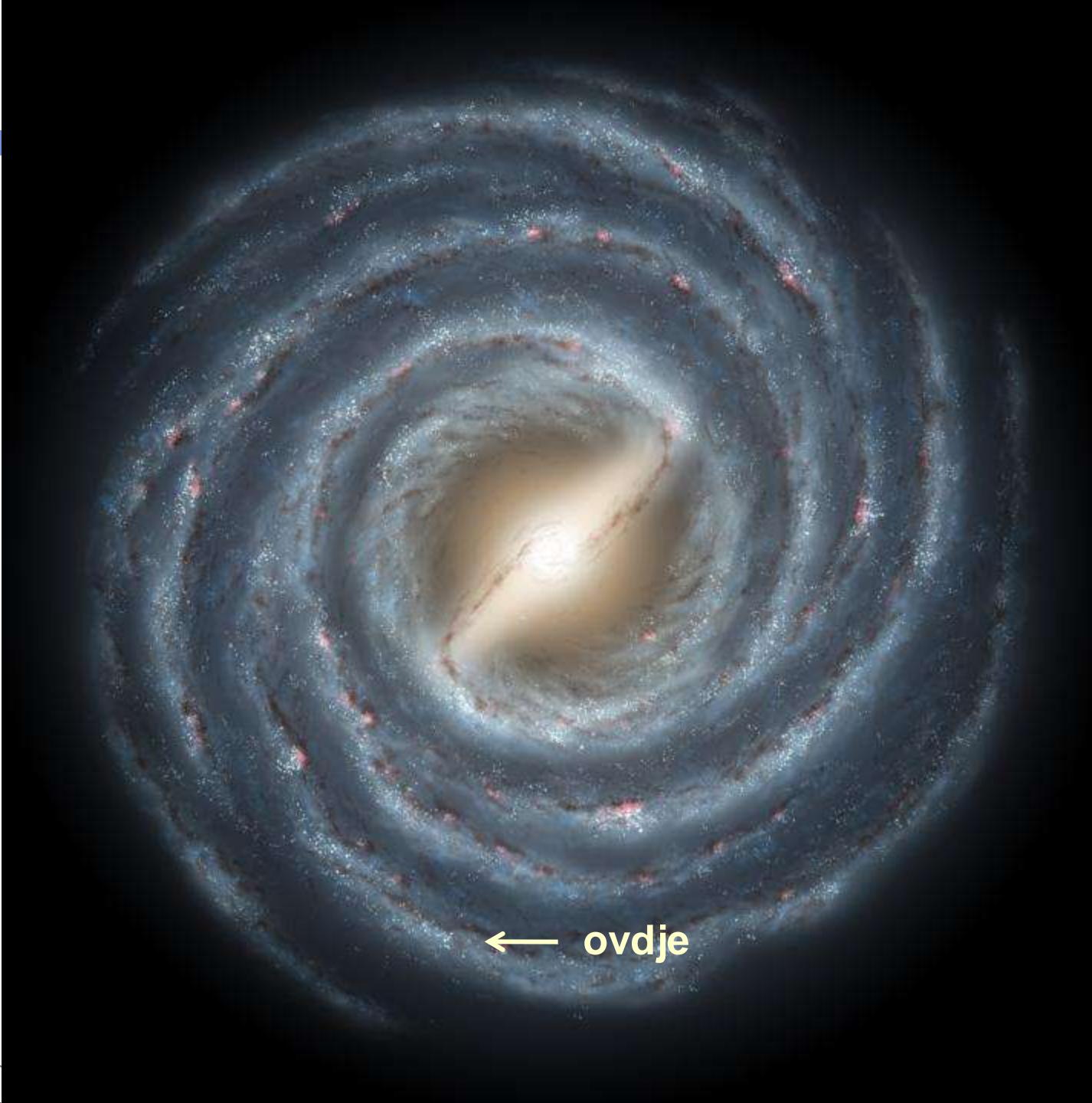
Dobriša Cesarić

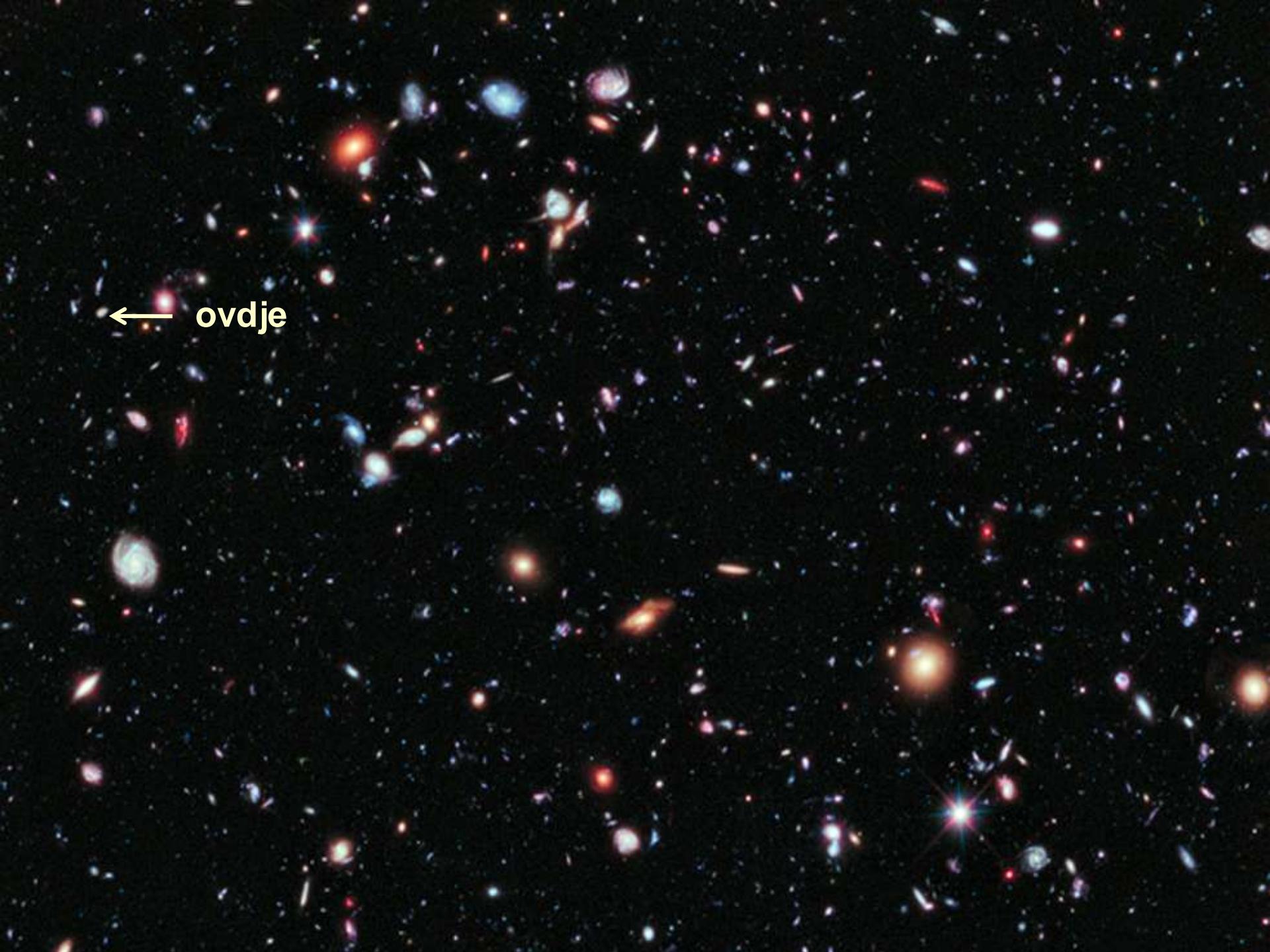
➤ Je li stvarno tako?



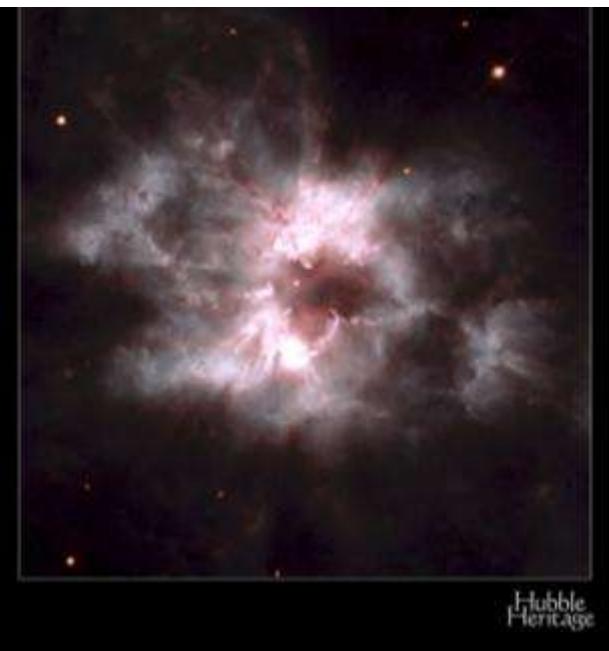
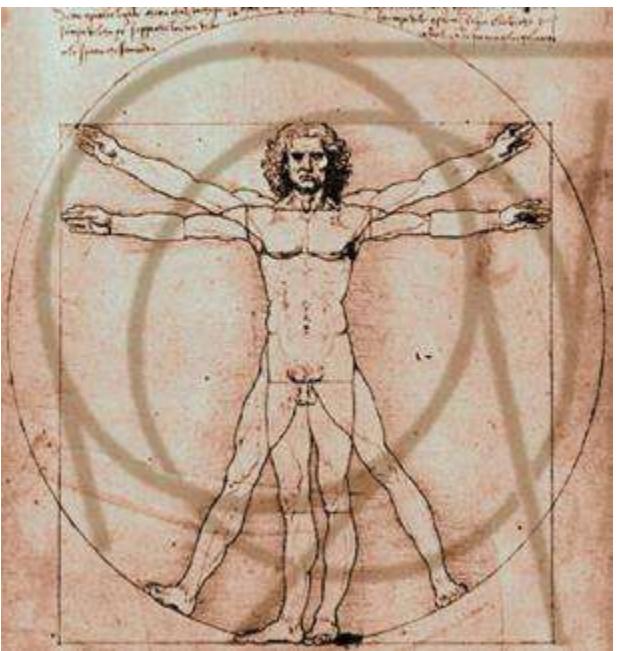
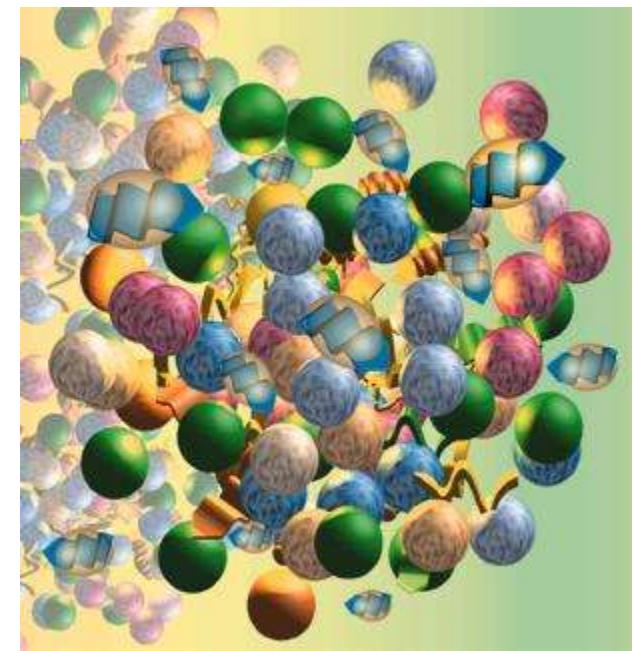
Earth as seen from Mars

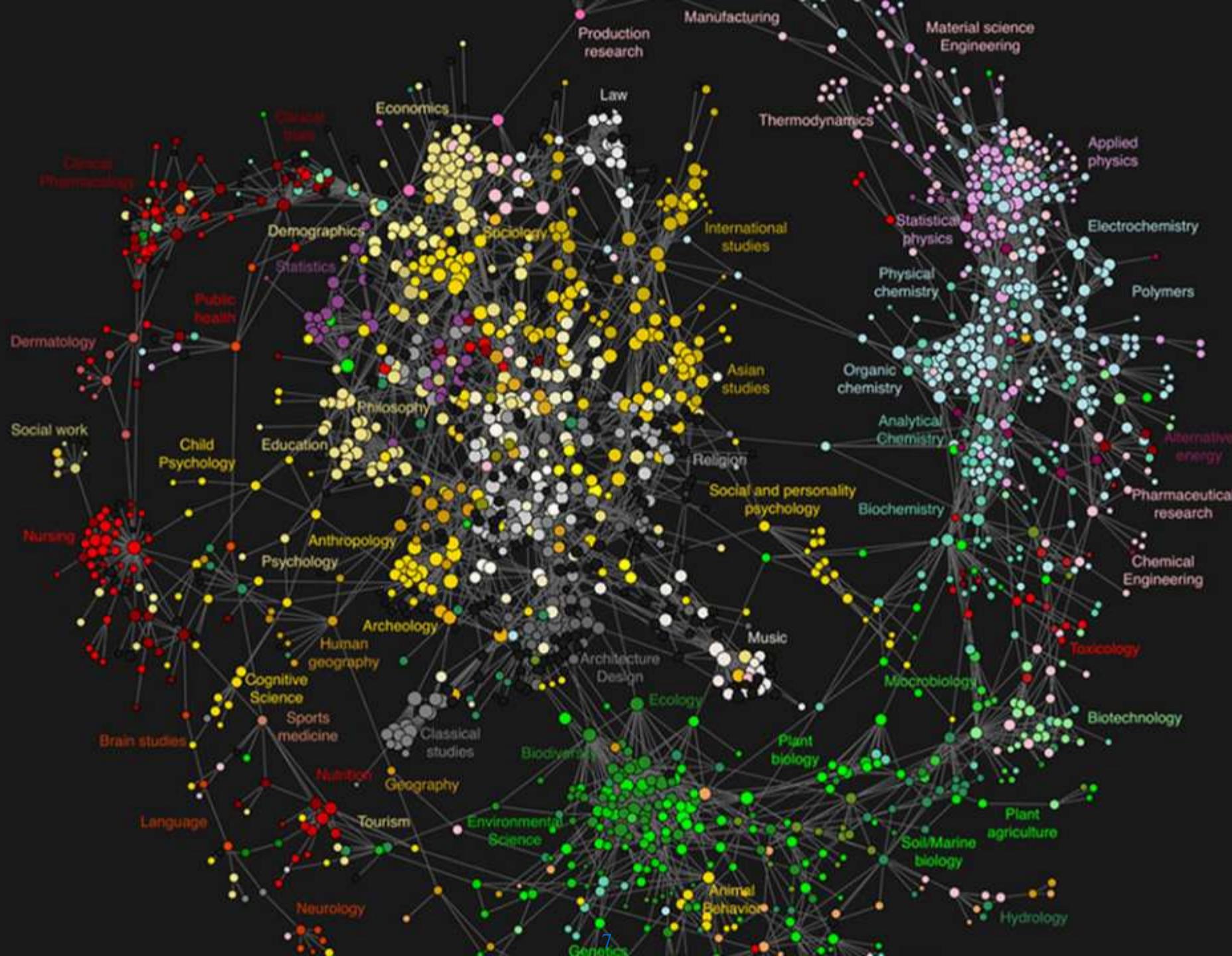






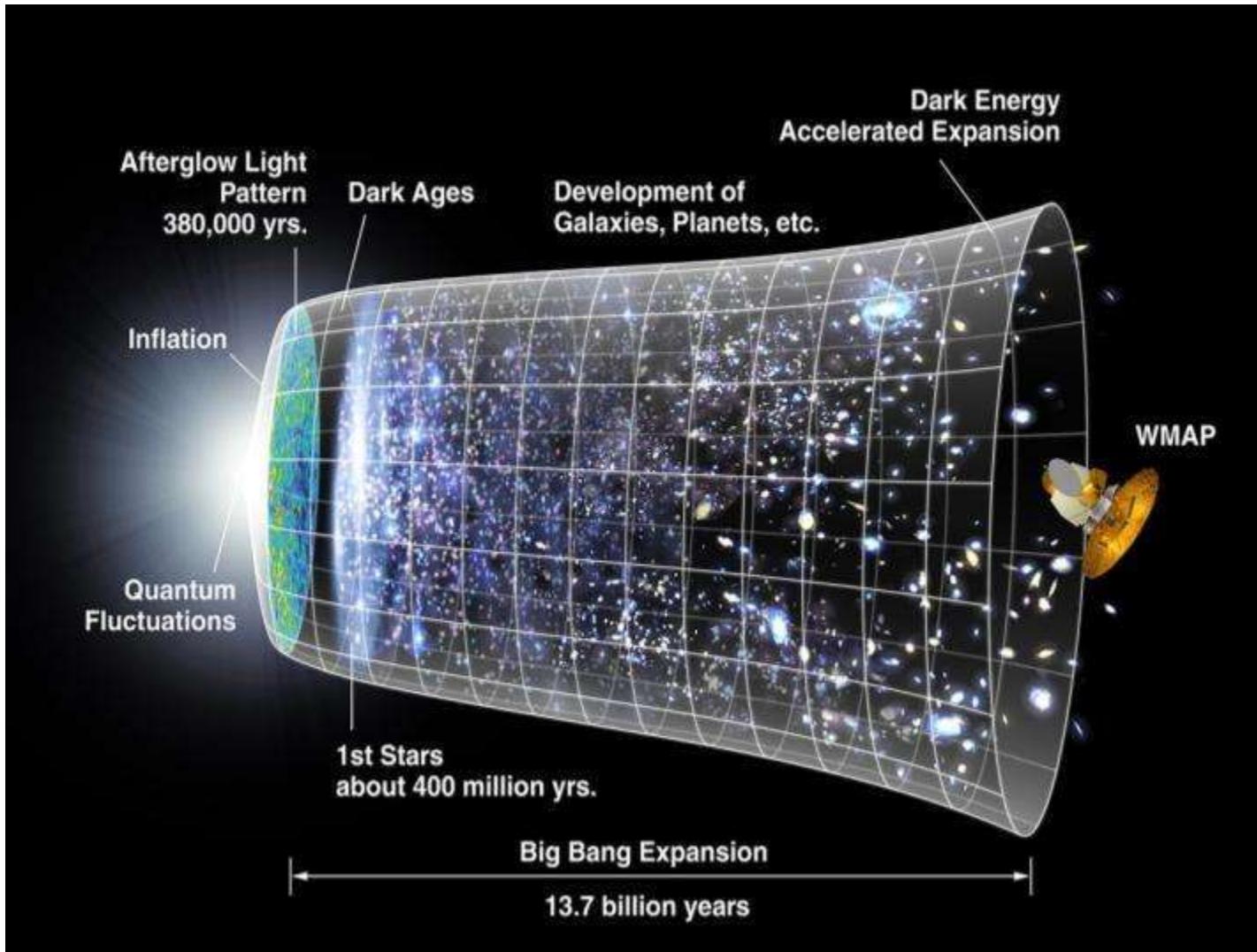
← ovdje





Što znamo?

➤ Kako se razvijao svemir



Što znamo?

➤ Standardni model čestica i njihovih interakcija

Leptoni		
Electric Charge		
Tau	-1	0
Tau Neutrino		
Mion	-1	0
Mion Neutrino		
Elektron	-1	0
Elektron Neutrino		

Jaka		
Gluoni (8)		
Kvarkovi		
Mezoni		
Barioni		
Jezgre		

Elektromagnetska		
Foton		
Atomi	Svetlost	Kemija
		Elektronika

Kvarkovi		
Električni naboј		
Dno	-1/3	2/3
Vrh		
Strani	-1/3	2/3
Šarmantni		
Dolje	-1/3	2/3
Gore		
svaki kvark R, B, G 3 boje		

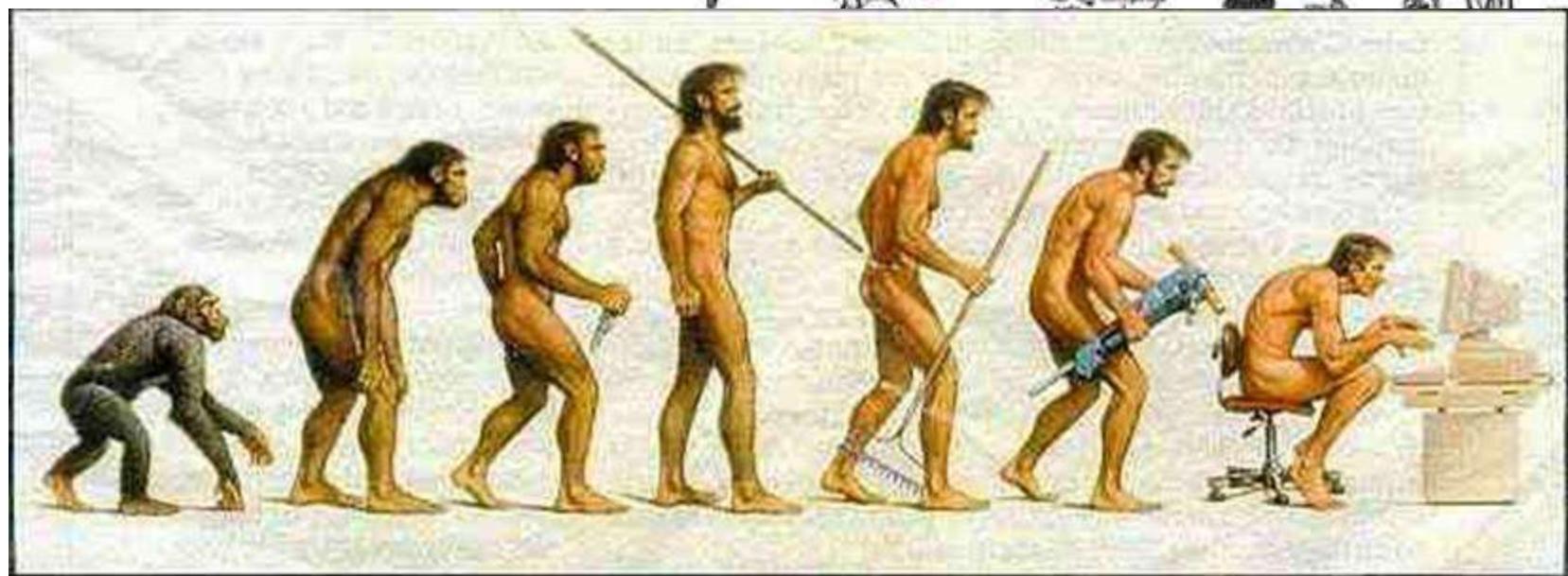
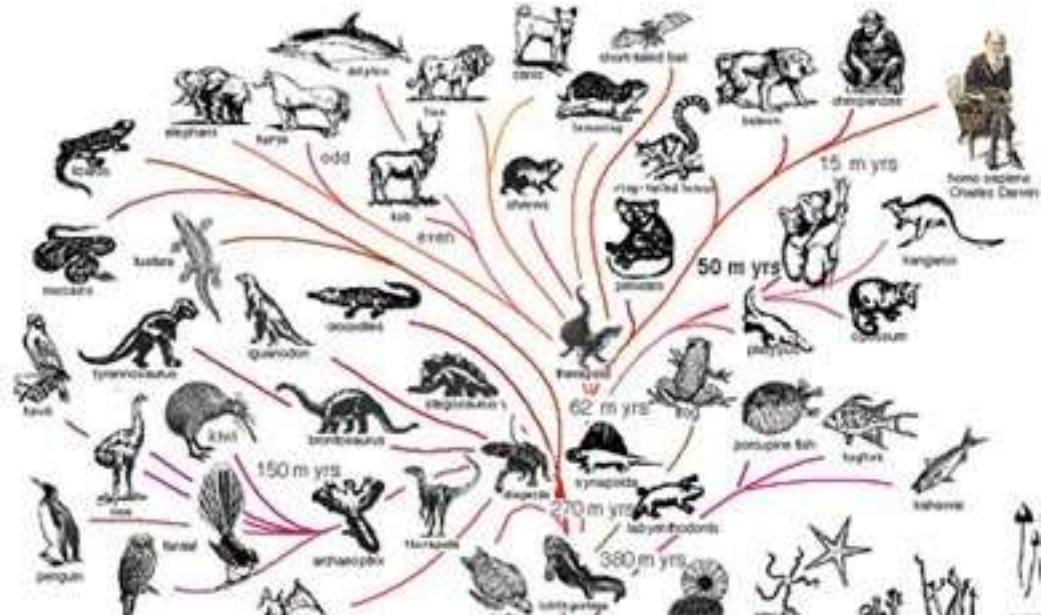
Gravitacija		
Graviton ?		
Sunčev sustav	Galaksije	Crne rupe

Slaba		
Bozoni (W,Z)		
Raspad neutrona	Beta radioaktivnost	Interakcije neutrina
		Izgaranja sunca

The particle drawings are simple artistic representations

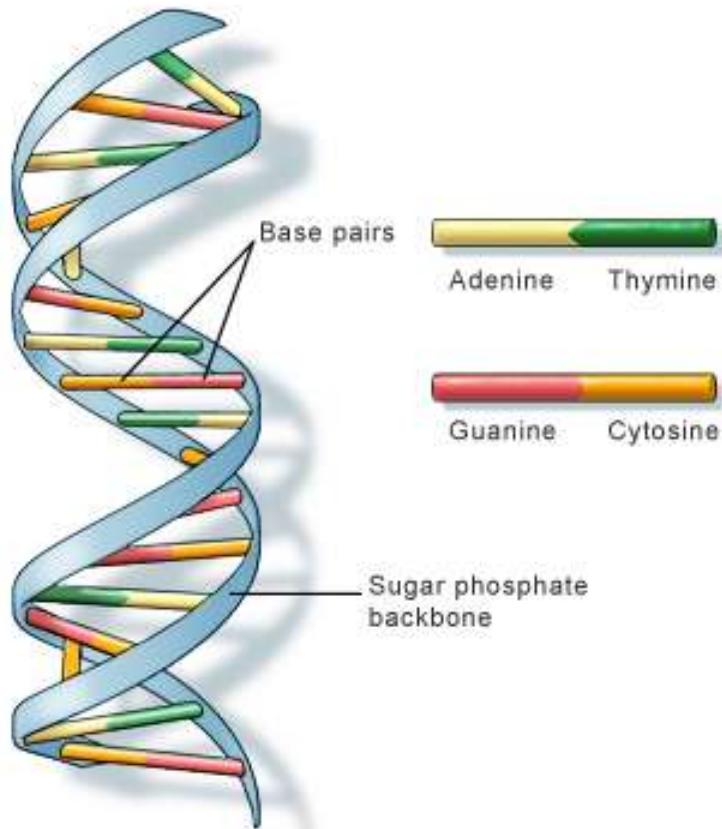
Što znamo?

- Darwinova teorija prirodne selekcije

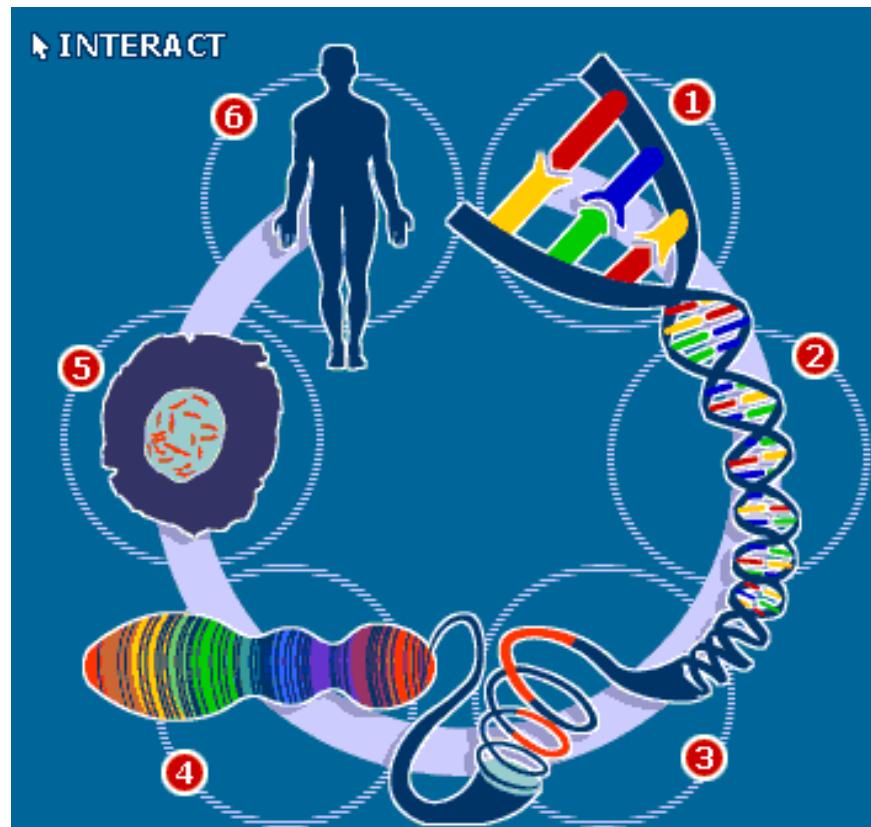


Što znamo?

➤ DNA



U.S. National Library of Medicine



Što još ne znamo?

- Kako je nastao i kako će završiti svemir?
- Od čega je izgrađen svemir?
- Kako mozak proizvodi svijest?
- Zašto ljudi imaju tako malo gena?
- Do koje mjere su povezane genetske varijacije i osobno zdravlje?
- Mogu li se zakoni fizike ujediniti?
- Koliko se može produljiti životni vijek ljudi?
- Kako se regeneriraju organi?
- Kako ćelija iz kože postane ćelija živca?
- Kako se iz oplođene jajne stanice razvije živo biće?
- Kako funkcioniра unutrašnjost Zemlje?
- Jesmo li sami u svemiru?
- Kako i kada se pojavio život na zemljji?
- Što određuje raznolikost vrsta?
- Koje genetske promjene su nas učinile ljudima?

Što još ne znamo?

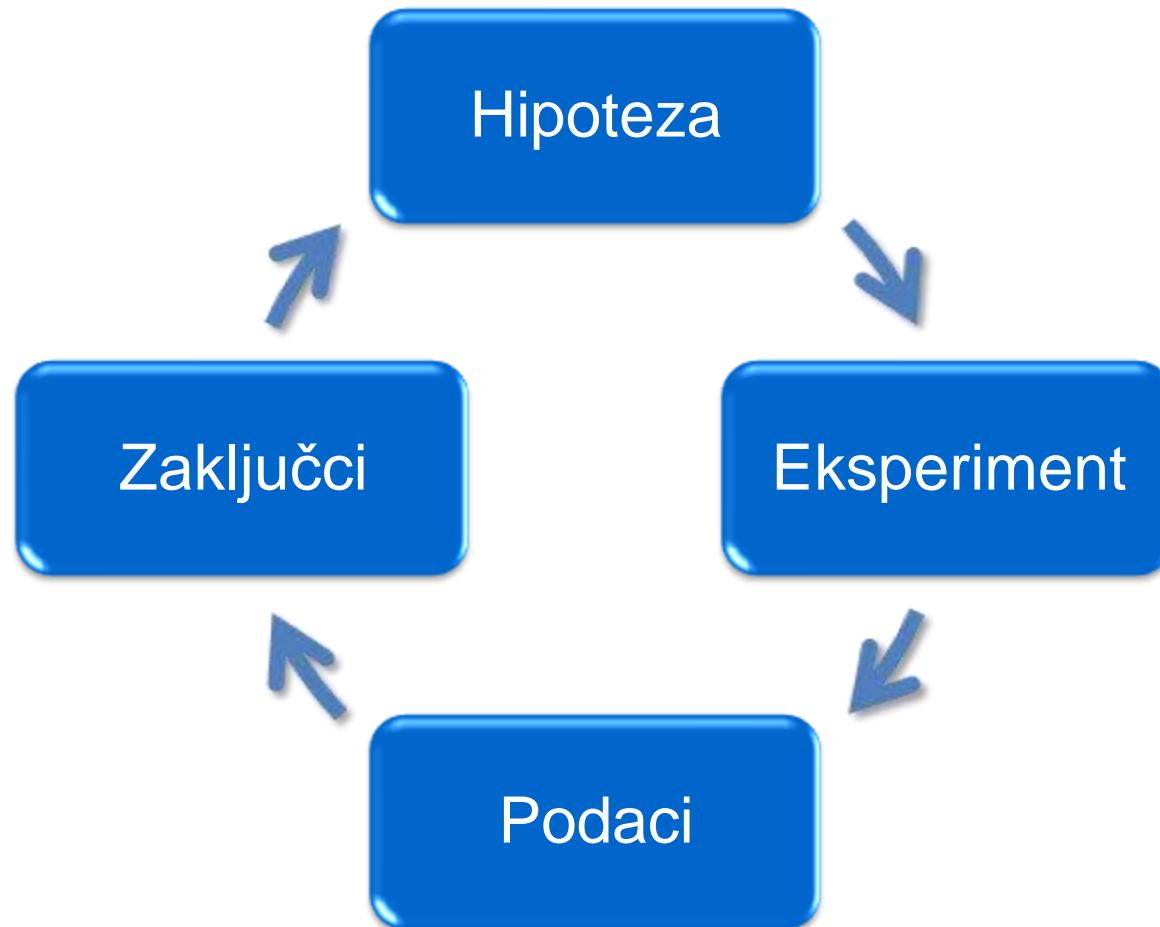
- Kako se pohranjuje i koristi pamćenje?
- Kako je evoluiralo kooperativno ponašanje?
- Koliko daleko možemo iskoristiti kemijsko samo-sastavljanje?
- Koje su granice konvencionalnog računanja?
- Možemo li selektivno isključiti imunološke odzive?
- Postoji li dublje razumijevanje kvantne neodređenosti i ne-lokalnosti?
- Je li u principu moguća efikasno cjepivo protiv HIV-a?
- Što može zamijeniti jeftinu naftu? I kada?
- Hoće li Malthus i dalje biti u krivu?
- Je li vrijeme iluzija?
- Zašto spavamo?
- Zašto funkcioniра placebo efekt?
- Koja je funkcija ne-kodirajućeg dijela DNA?
- Hoće li šume usporiti globalno zagrijavanje? Ili ga ubrzati?

Što još ne znamo?

- Što se događa s informacijom u crnoj rupi?
- Što uzrokuje 'ledena doba'?
- Kako mozak računa pokrete?
- Zašto se Zemljini polovi izmjenjuju?
- Kako je evoluirao ljudski jezik?
- Zašto ne možemo predvidjeti vrijeme?
- Zašto ne razumijemo turbulencije?
- Je li svemir u biti izgrađen od informacija?
- Zašto se neke bolesti pretvore u epidemije?
- Zašto umiremo?
- Što uzrokuje gravitaciju?
- Zašto ne mogu ponovo izrasti izgubljeni udovi?
- Zašto čestice imaju masu? I to različitu!

U što smo absolutno uvjereni?

- Da će se odgovori na ova pitanja, prije ili kasnije, pronaći isključivo
ZNANSTVENOM METODOLOGIJOM





SO YOU'RE TELLING ME



PEOPLE FROM YOUR COUNTRY THINK
THAT FACEBOOK GIVES US FOOD FOR
EVERY LIKE?

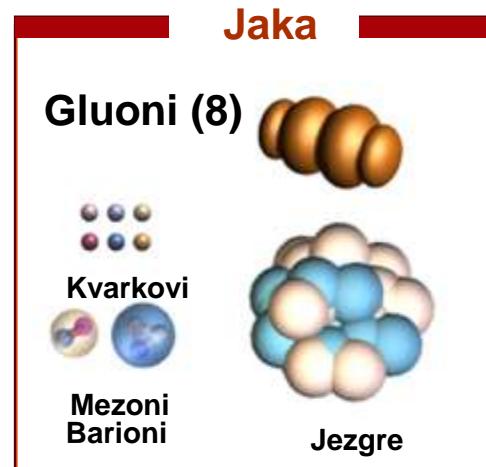


Od čega je izgrađen svemir?

Što znamo?

➤ Standardni model čestica i njihovih interakcija

Leptoni		
Electric Charge		
Tau	-1	0
Tau Neutrino		
Mion	-1	0
Mion Neutrino		
Elektron	-1	0
Elektron Neutrino		



Kvarkovi

Kvarkovi		
Električni naboj		
Dno	-1/3	2/3
Vrh		
Strani	-1/3	2/3
Šarmantni		
Dolje	-1/3	2/3
Gore		
svaki kvark R, B, G 3 boje		



The particle drawings are simple artistic representations

Čestice

Antičestice

Leptoni

Electric Charge		
Tau	-1	0
Mion	-1	0
Elektron	-1	0

Tau Neutrino
Mion Neutrino
Elektron Neutrino

Antileptoni

Electric Charge		
Antitau	1	0
Antimion	1	0
Pozitron	1	0

Antitau Neutrino
Antimion Neutrino
Antielektron Neutrino

Kvarkovi

Električni naboј		
Dno	-1/3	2/3
Strani	-1/3	2/3
Dolje	-1/3	2/3

Vrh
Šarmantni
Gore

svaki kvark:**R**, **B**, **G** 3 boje

Antikvarkovi

Električni naboј		
Antidno	1/3	-2/3
Antistrani	1/3	-2/3
Antidolje	1/3	-2/3

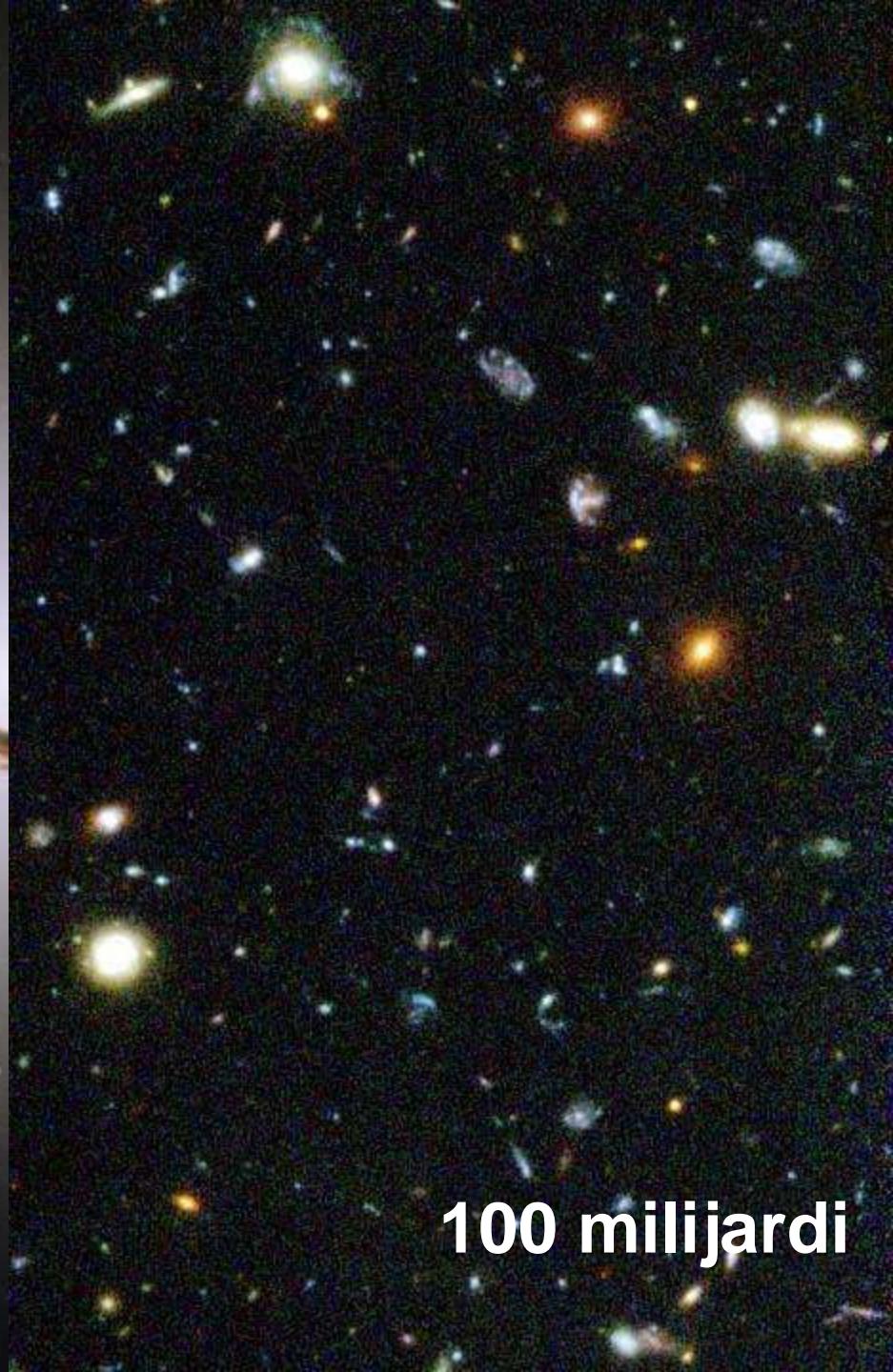
Antivrh
Antišarmantni
Antigore

svaki antikvark:**R**, **B**, **G** 3 antiboje

100 miliardi



100 miliardi



“Normalna tvar”

4%

Tamna energija

74%

Tamna tvar

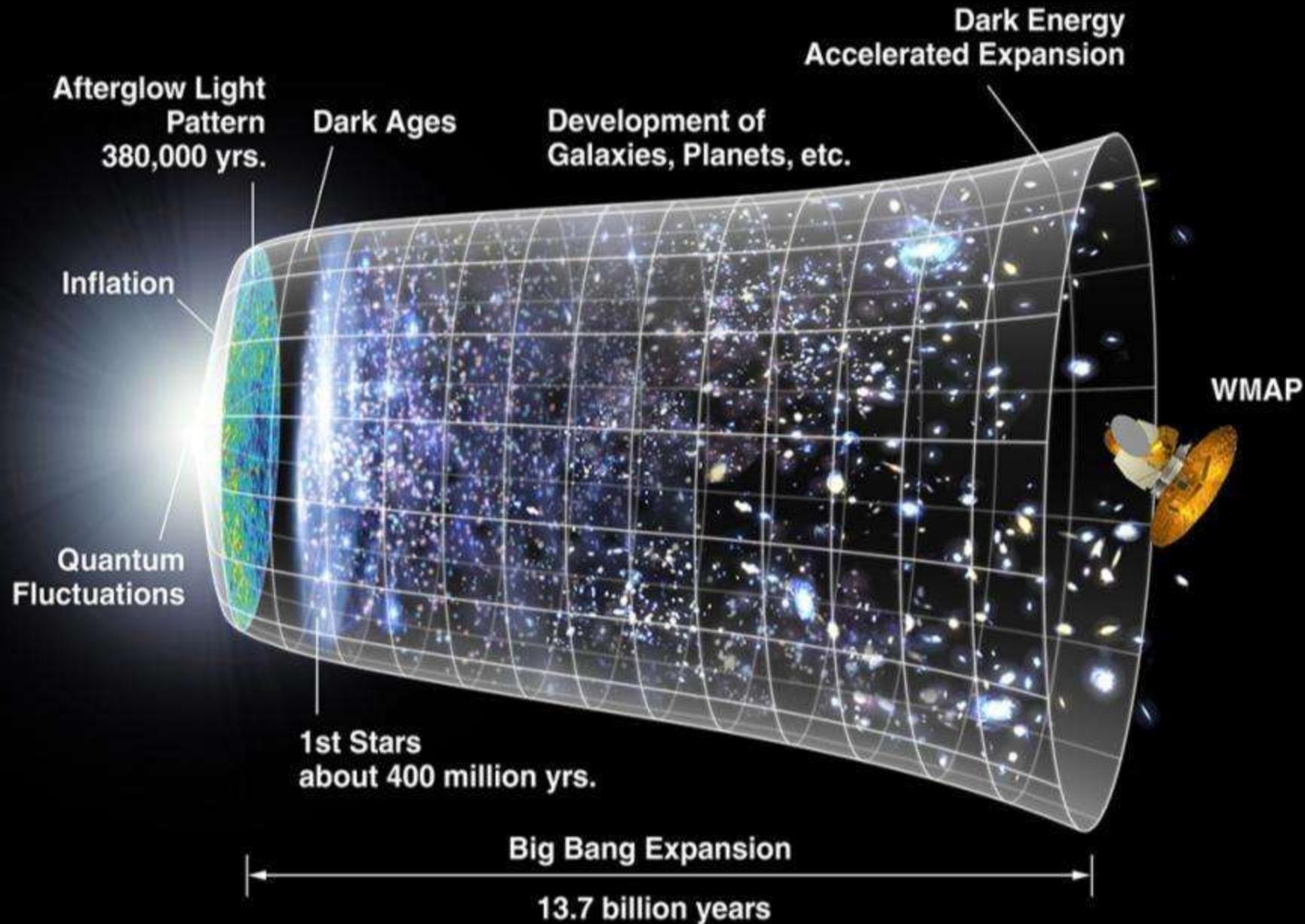
22%



The background of the image is a dark, textured space scene featuring numerous small white stars of varying brightness. A prominent, bright, diffuse band of light, characteristic of a spiral galaxy's disk, stretches horizontally across the center. The edges of this band appear slightly darker and more defined. In the upper right quadrant, there is a distinct, smaller, brighter cluster of stars, possibly a satellite galaxy or a dense star-forming region.

Kako to znamo?

Od Velikog praska do današnjeg dana



10^{-43} s Era kvantne gravitacije



**Gravitacija se odvaja
kao posebna sila,
ostale sile ostaju
ujedinjene u jednu
(Veliko ujedinjenje)**

$t < 10^{-43}$ s : The Big Bang

The universe is considered to have expanded from a single point with an infinitely high energy density (infinite temperature). Is there a meaning to the question what existed before the big bang?

$t = 10^{-43}$ s, 10^{32} K (10^{19} GeV, 10^{-34} m) : Gravity "freezes" out

All particle types (quarks, leptons, gauge bosons, and undiscovered particles e.g. Higgs, sparticles, gravitons) and their anti-particles are in a thermal equilibrium (being created and annihilated at equal rate). These coexist with photons (radiation).

Through a phase transition gravity "froze" out and became distinct in its action from the weak, electromagnetic and strong forces. The other three forces could not be distinguished from one another in their action on quarks and leptons. This is the first instance of the breaking of symmetry amongst the forces.

10^{-35} s Era velikog ujedinjenja



Inflacija prestaje, širenje se nastavlja. Prestaje era velikog ujedinjenja. Jaka i elektroslaba sila počinju se razlikovati.

t - 10^{-35} s, 10^7 K (10^6 GeV, 10^{32} m) : Inflation

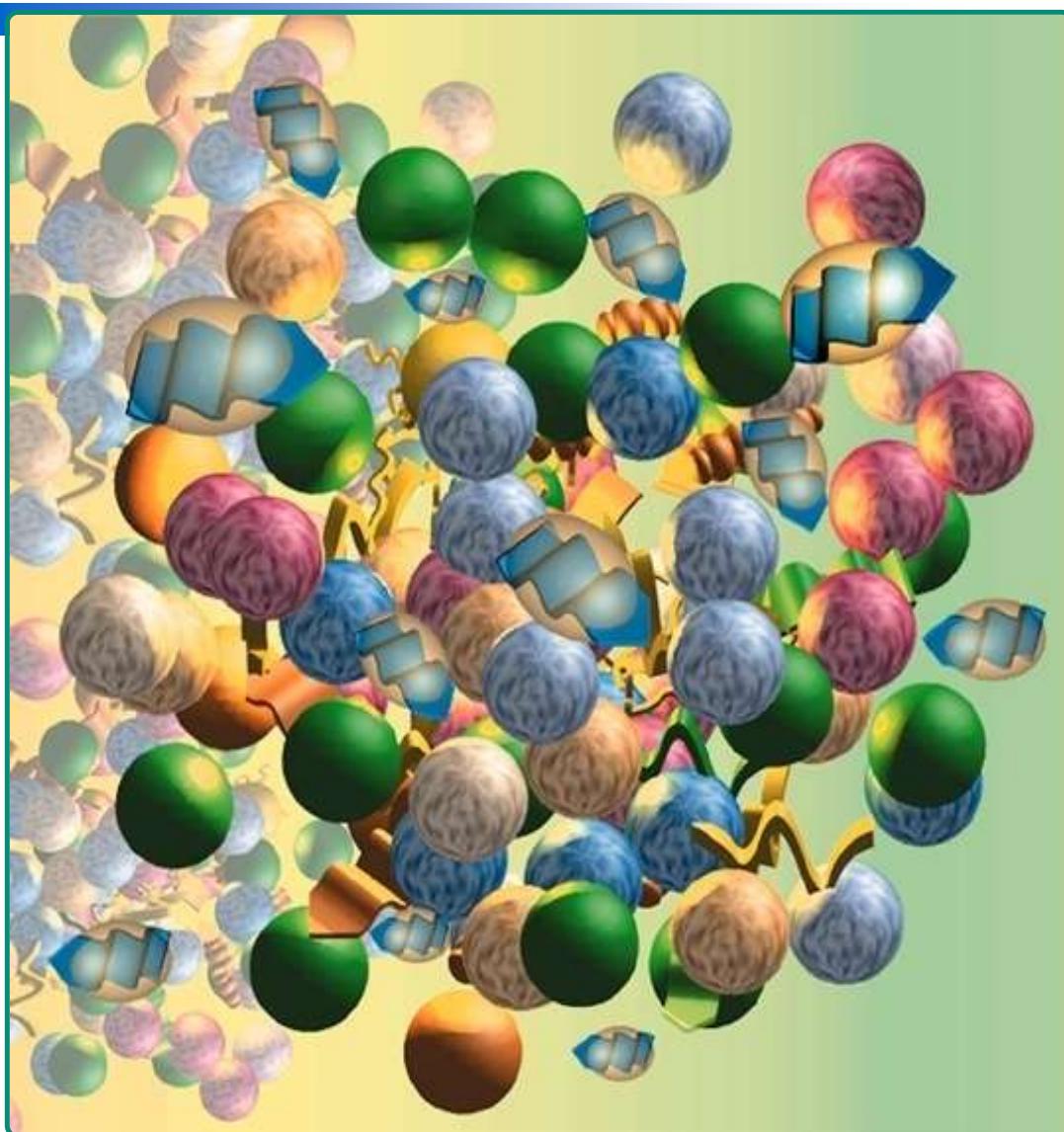
The rate of expansion increases exponentially for a short period. The universe doubled in size every 10^{-35} s. Inflation stopped at around 10^{-32} s. The universe increased in size by a factor of 10^{10} . This is equivalent to an object the size of a proton swelling to 10^{10} light years across.

However the presently visible universe was only 3 m in size after inflation. This solves the problems of ‘horizon’ (how is it possible for two opposing parts of the present universe to be at the same temperature when they cannot have interacted with each other before recombination) and ‘flatness’ (density of matter is close to the critical density).

t - 10^{-32} s : Strong forces freezes out

Through another phase transition the strong force “freezes” out and a slight excess of matter over anti-matter develops. This excess, at a level of 1 part in a billion, is sufficient to give the presently observed predominance of matter over anti-matter. The temperature is too high for quarks to remain clumped to form neutrons or protons and so exist in the form of a quark gluon plasma. The LHC can study this by colliding together high energy nuclei.

10^{-10} s Elektroslaba era

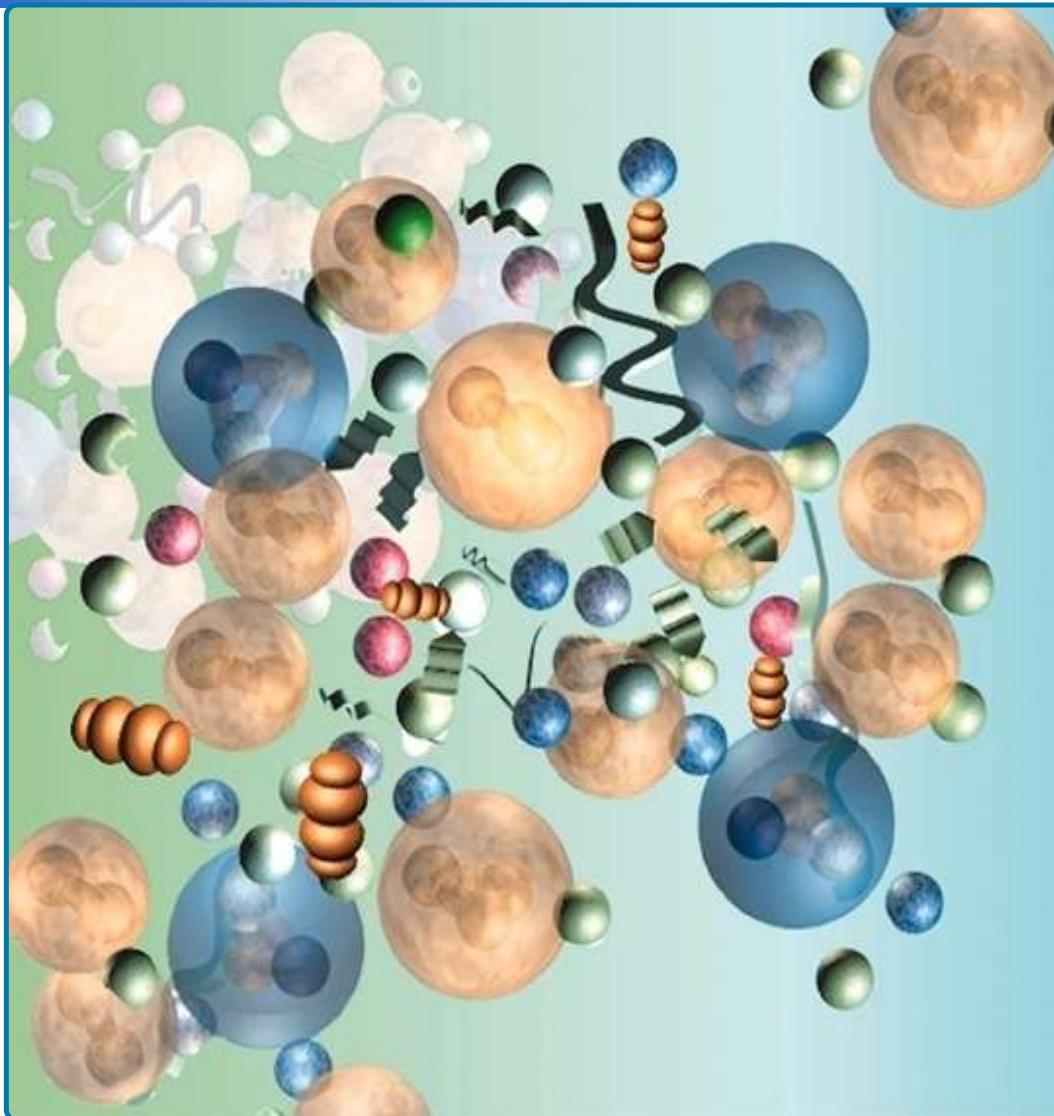


Razdvaja se elektroslaba sila

**$t = 10^{-10}$ s, 10^{15} K (100 GeV, 10^{18} m) :
Electromagnetic and Weak Forces
separate**

The energy density corresponds to that at LEP. As the temperature fell the weak force "freezes" out and all four forces become distinct in their actions. The antiquarks annihilate with the quarks leaving a residual excess of matter. W and Z bosons decay. In general unstable massive particles disappear when the temperature falls to a value at which photons from the black-body radiation do not have sufficient energy to create a particle-antiparticle pair.

10^{-4} s Stvaranje protona i neutrona



Kvarkovi se kombiniraju i ujedinjavaju u protone i neutrone

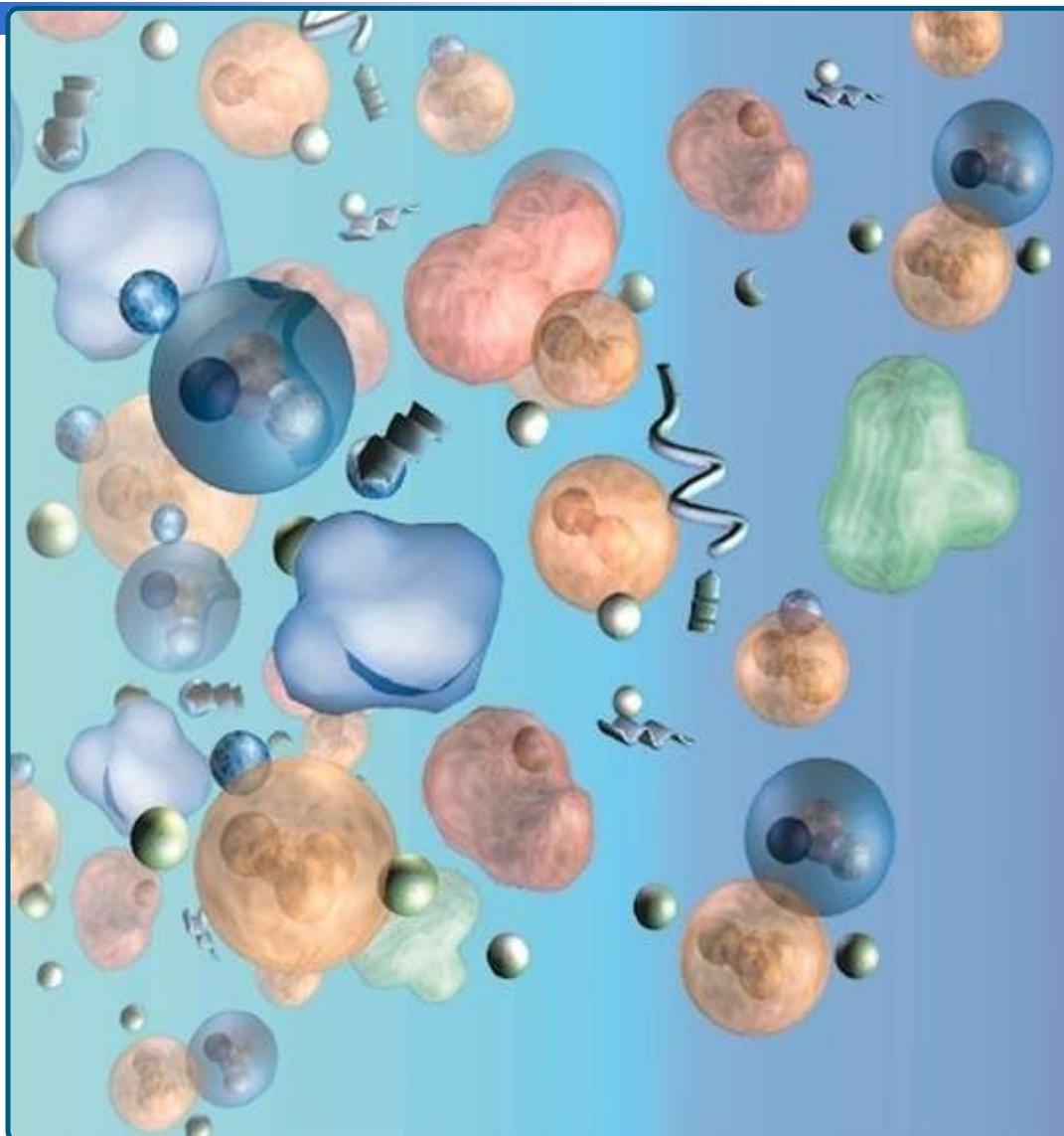
**$t = 10^{-4}$ s, 10^{13} K (1 GeV, 10^{16} m) :
Protons and Neutrons form**

The universe has grown to the size of our solar system. As the temperature drops quark-antiquark annihilation stops and the remaining quarks combine to make protons and neutrons.

**$t = 1$ s, 10^{10} K (1 MeV, 10^{15} m) :
Neutrinos decouple**

The neutrinos become inactive (essentially do not participate further in interactions). The electrons and positrons annihilate and are not recreated. An excess of electrons is left. The neutron-proton ratio shifts from 50:50 to 25:75.

100 s

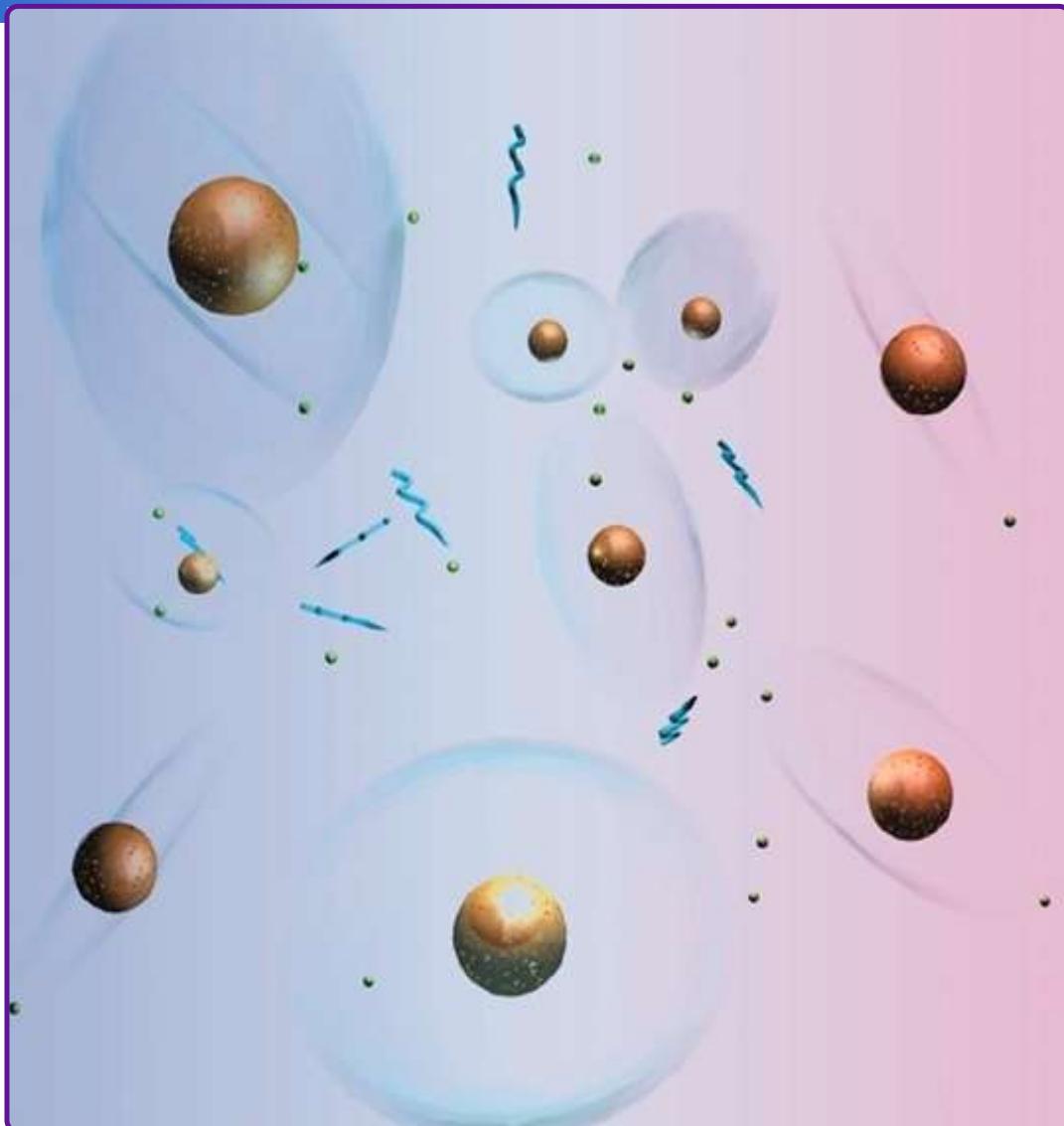


Protoni i neutroni se kombiniraju u jezgru helij

$t = 3 \text{ minutes, } 10^9 \text{ K (0.1 MeV, } 10^{12} \text{ m)}$: Nuclei are formed

The temperature is low enough to allow nuclei to be formed. Conditions are similar to those that exist in stars today or in thermonuclear bombs. Heavier nuclei such as deuterium, helium and lithium soak up the neutrons that are present. Any remaining neutrons decay with a time constant of ~ 1000 seconds. The neutron-proton ratio is now 13:87. The bulk constitution of the universe is now in place consisting essentially of protons (75%) and helium nuclei. The temperature is still too high to form any atom and electrons form a gas of free particles.

300000 years Atomi i era svjetla



Svemir postaje proziran
i ispunjen svjetлом

**$t = 300\ 000 \text{ years}, 6000 \text{ K} (0.5 \text{ eV}, 10^9 \text{ m})$:
Atoms are created**

Electrons begin to stick to nuclei. Atoms of hydrogen, helium and lithium are created. Radiation is no longer energetic enough to break atoms. The universe becomes transparent. Matter density dominates. Astronomy can study the evolution of the Universe back to this time.

1000 M years

Stvaranje galaksija

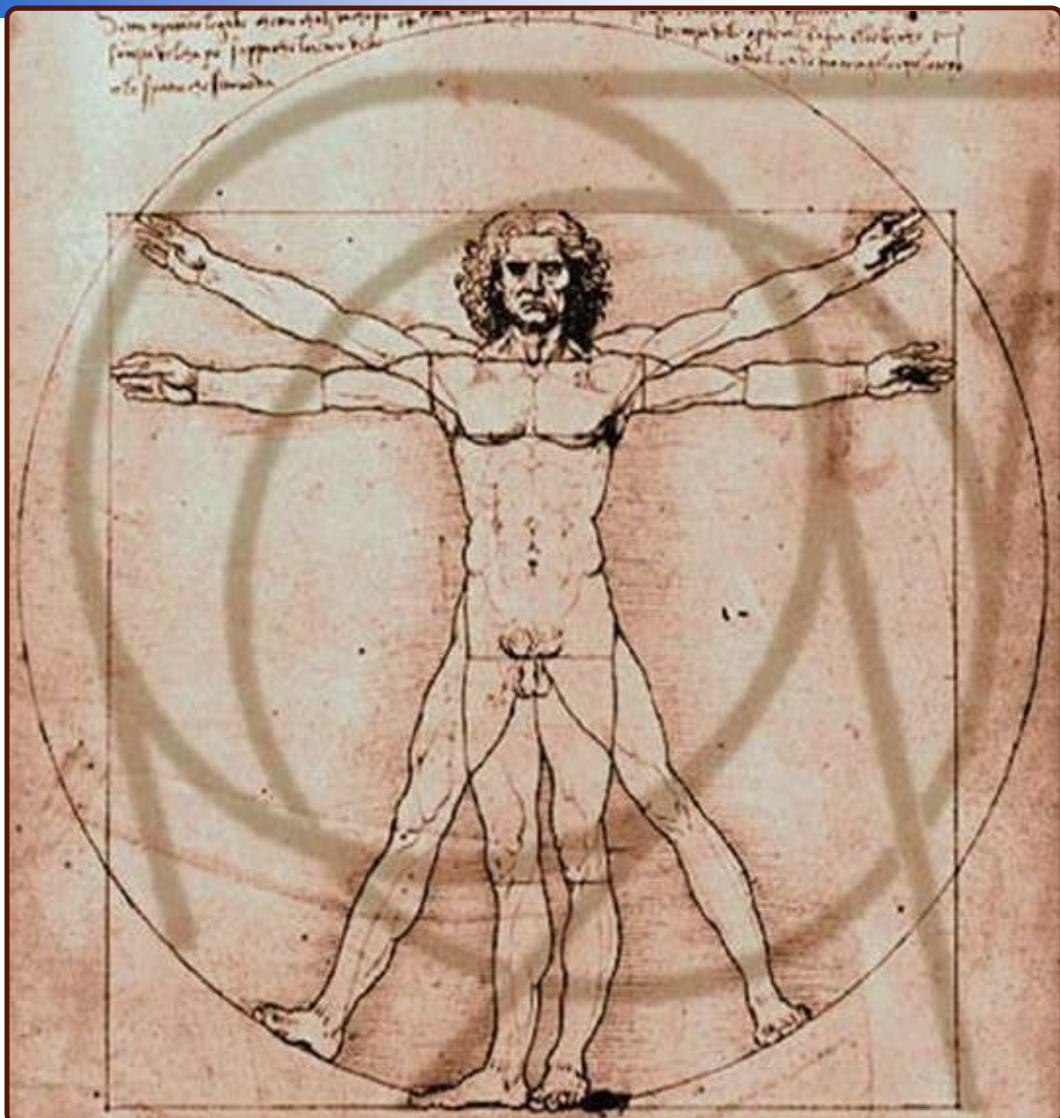


Galaksije se počinju stvarati

$t = 10^9$ years, 18 K : Galaxy Formation

Local mass density fluctuations act as seeds for stellar and galaxy formation. The exact mechanism is still not understood. Nucleosynthesis, synthesis of heavier nuclei such as carbon up to iron, starts occurring in the thermonuclear reactors that are stars. Even heavier elements are synthesized and dispersed in the brief moment during which stellar collapse and supernovae explosions occur.

15000 M years Danas

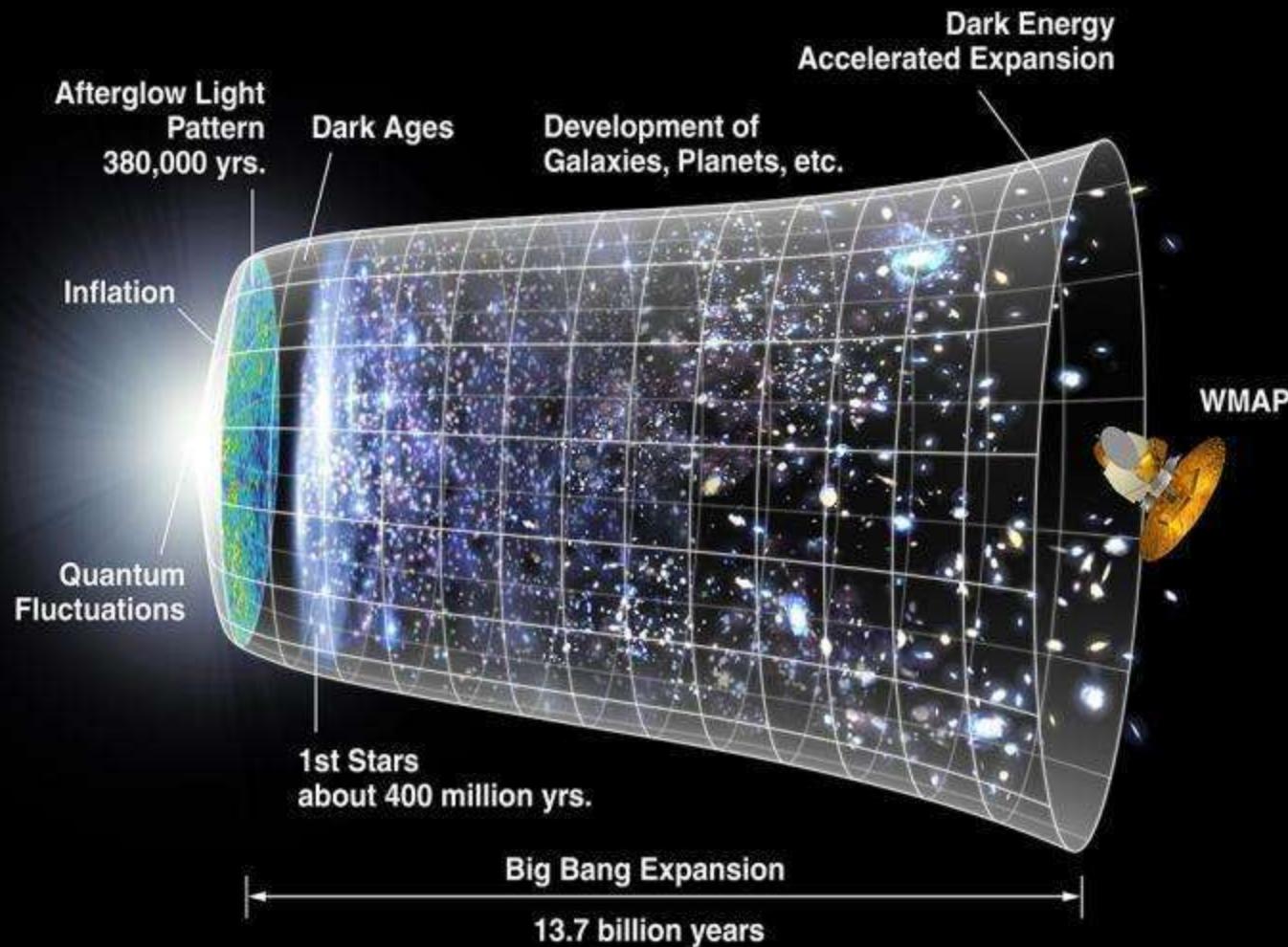


Čovjek se počeo pitati
odakle sve ovo!

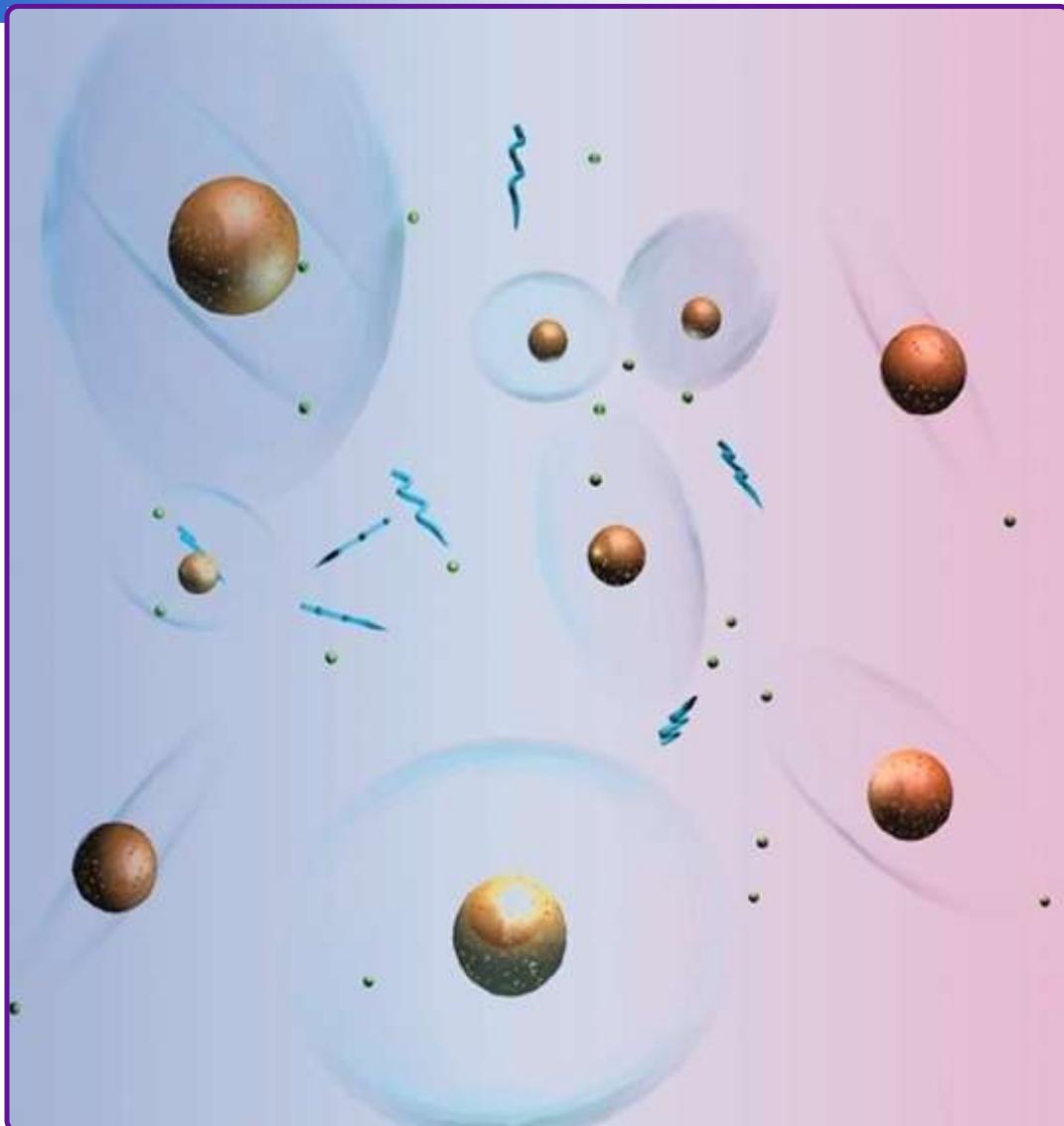
$t = 15 \times 10^9$ years, 3 K : Humans

The present day. Chemical processes have linked atoms to form molecules. From the dust of stars and through coded messages (DNA) humans emerge to observe the universe around them.

Od Velikog praska do današnjeg dana



300000 years Atomi i era svjetla



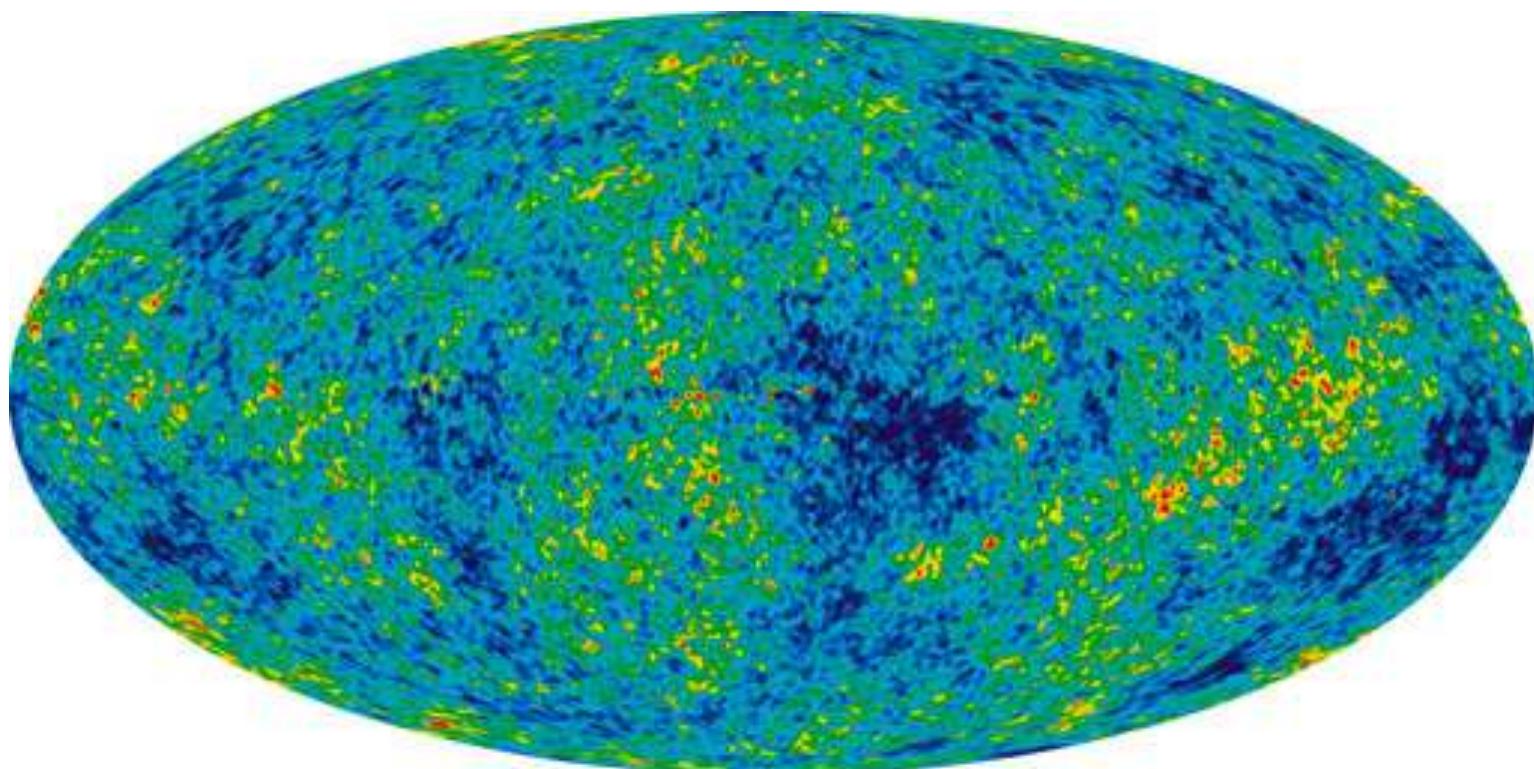
Svemir postaje proziran
i ispunjen svjetлом

**$t = 300\ 000 \text{ years}, 6000 \text{ K} (0.5 \text{ eV}, 10^9 \text{ m})$:
Atoms are created**

Electrons begin to stick to nuclei. Atoms of hydrogen, helium and lithium are created. Radiation is no longer energetic enough to break atoms. The universe becomes transparent. Matter density dominates. Astronomy can study the evolution of the Universe back to this time.

Wilkinson Microwave Anisotropy Probe

WMAP



Čestice

Antičestice

Leptoni

	Electric Charge		
Tau	-1	0	Tau Neutrino
Mion	-1	0	Mion Neutrino
Elektron	-1	0	Elektron Neutrino

Antileptoni

	Electric Charge		
Antitau	1	0	Antitau Neutrino
Antimion	1	0	Antimion Neutrino
Pozitron	1	0	Antielektron Neutrino

Kvarkovi

	Električni naboј			
Dno	-1/3	2/3		Vrh
Strani	-1/3	2/3		Šarmantni
Dolje	-1/3	2/3		Gore

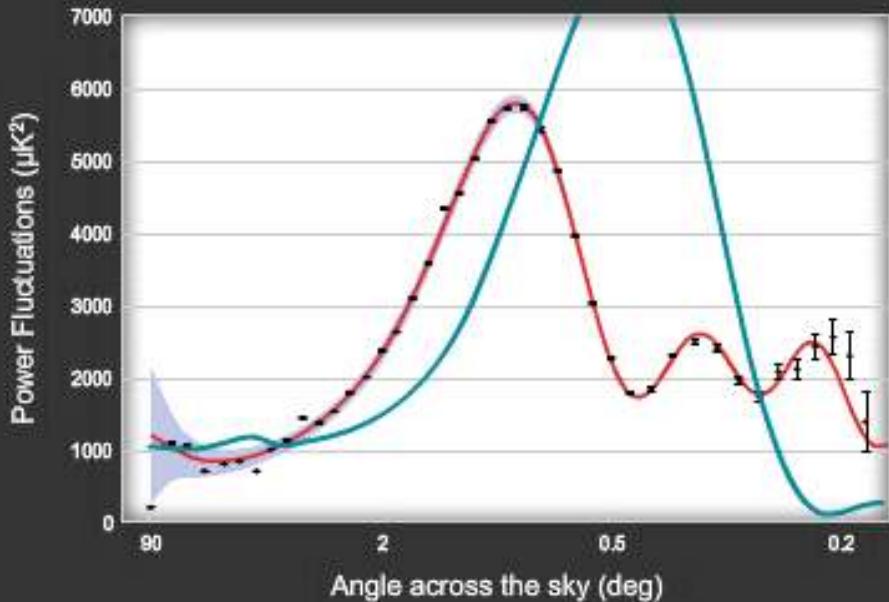
svaki kvark:**R**, **B**, **G** 3 boje

Antikvarkovi

	Električni naboј			
Antidno	1/3	-2/3		Antivrh
Antistrani	1/3	-2/3		Antišarmantni
Antidolje	1/3	-2/3		Antigore

svaki antikvark:**R**, **B**, **G** 3 antiboje

WMAP CMB Analyzer



Age: 9.1 billion years

Flatness: 1.00

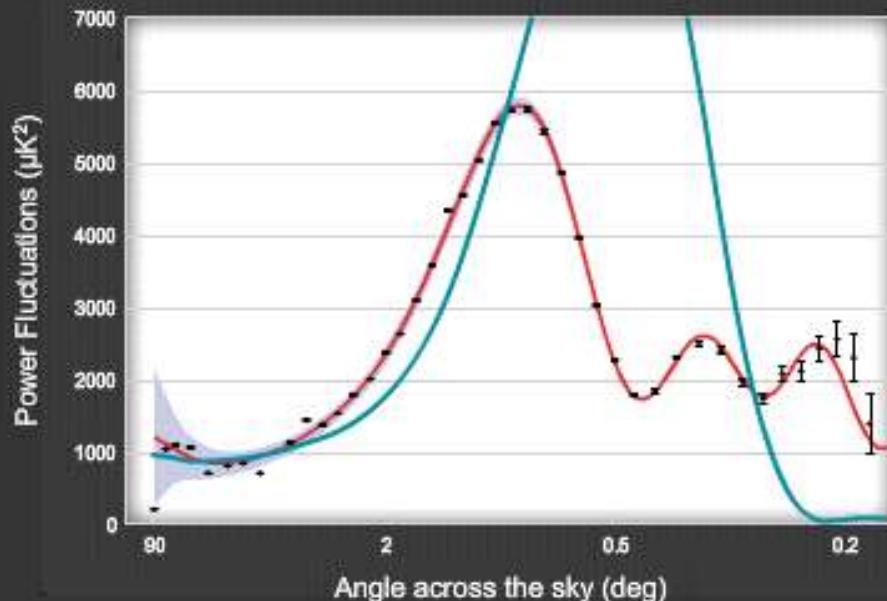
Power Spectrum Plot: This plot shows how temperature varies with the angular size of patches on the sky. This reveals the energy emitted by different size ripples of sound traveling through the early universe.

- Red line = analyzed sky / universe signal.
- Blue line = your simulated sky / universe signal.
- Black points with error bars = 'binned' (grouped) data to analyze data accuracy.
- Light blue area = likelihood of results being caused by random chance- only a concern at large scale (left).

ANSWER

RESET

WMAP CMB Analyzer



Age: 9.2 billion years

Flatness: 1.00

Pie Chart: Graphically shows the composition of your universe. The wedges compare the amount of each component; the size of the pie compares the total composition (matter + dark matter + dark energy) with the critical density (black circle).

- A universe at critical density is geometrically flat and probably infinite.
- A universe can have more or less than the critical density.
- Flatness - the term we use for closeness to critical density.

ANSWER

RESET

WMAP CMB Analyzer



Universe Content



Atoms



Cold Dark Matter



Dark Energy

Additional Properties

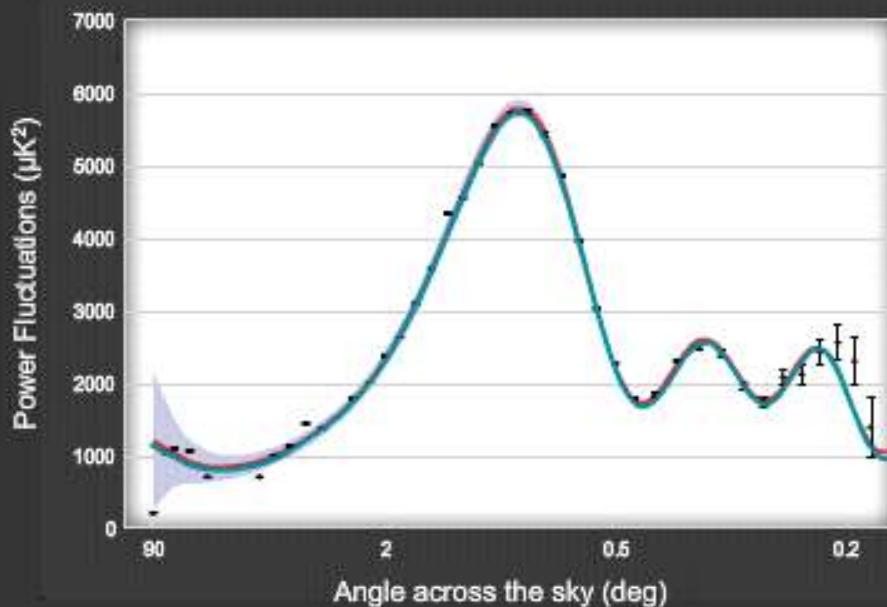
Hubble Constant



Reionization redshift



Spectral Index



Age: 13.7 billion years

Flatness: 1.00

Pie Chart: Graphically shows the composition of your universe. The wedges compare the amount of each component; the size of the pie compares the total composition (matter + dark matter + dark energy) with the critical density (black circle).

- A universe at critical density is geometrically flat and probably infinite.
- A universe can have more or less than the critical density.
- Flatness - the term we use for closeness to critical density.

ANSWER

RESET

“Normalna tvar”
4%



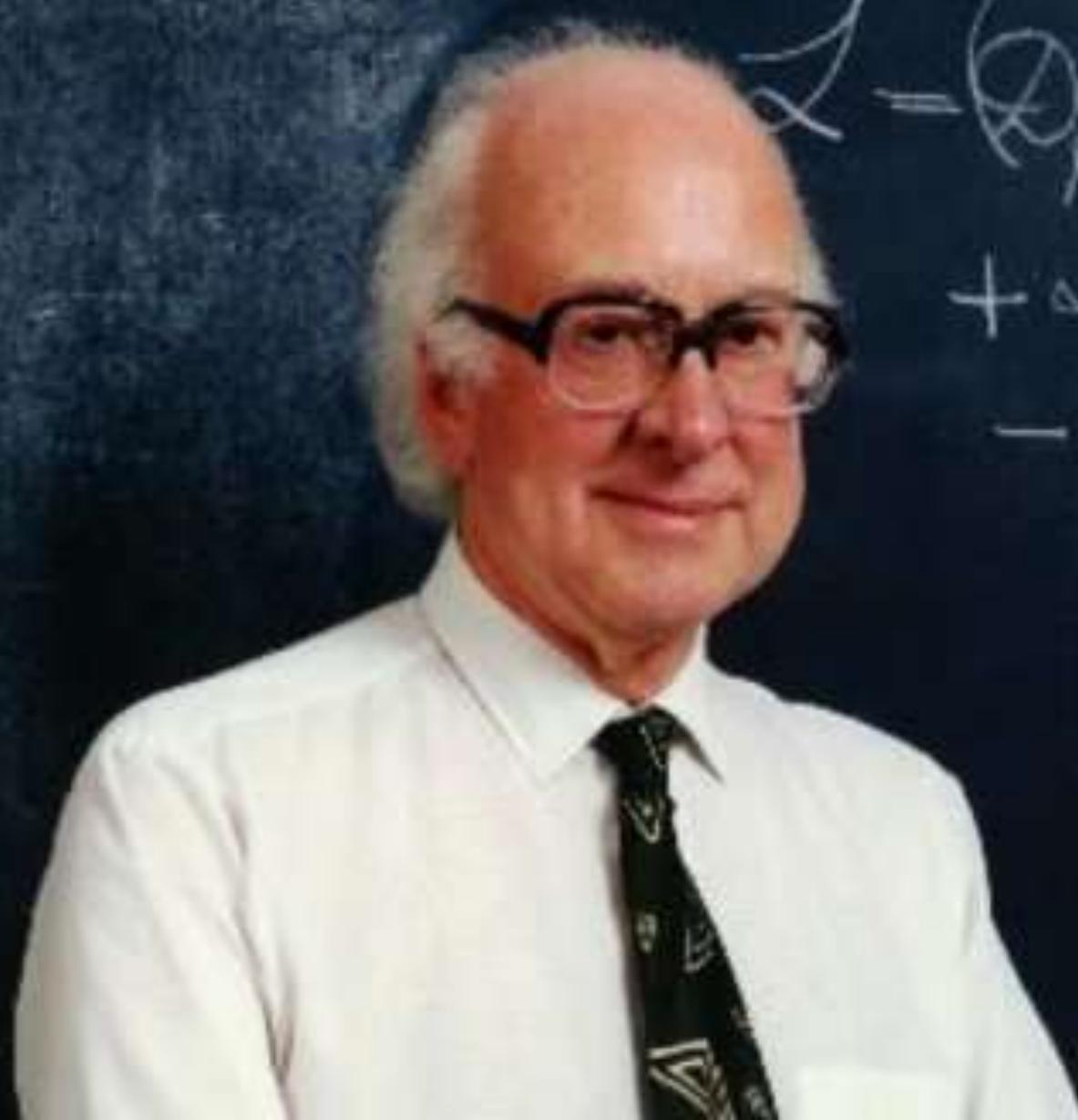
100 milijardi

SVE OVO JE SAMO 4% SVEMIRA!

100 milijardi



Zašto čestice imaju masu?



Peter
Higgs

Čestice

Antičestice

Leptoni

Electric Charge		
Tau	-1	0
Mion	-1	0
Elektron	-1	0

Tau Neutrino
Mion Neutrino
Elektron Neutrino

Antileptoni

Electric Charge		
Antitau	1	0
Antimion	1	0
Pozitron	1	0

Antitau Neutrino
Antimion Neutrino
Antielektron Neutrino

Kvarkovi

Električni naboј		
Dno	-1/3	2/3
Strani	-1/3	2/3
Dolje	-1/3	2/3

Vrh
Šarmantni
Gore

svaki kvark:**R**, **B**, **G** 3 boje

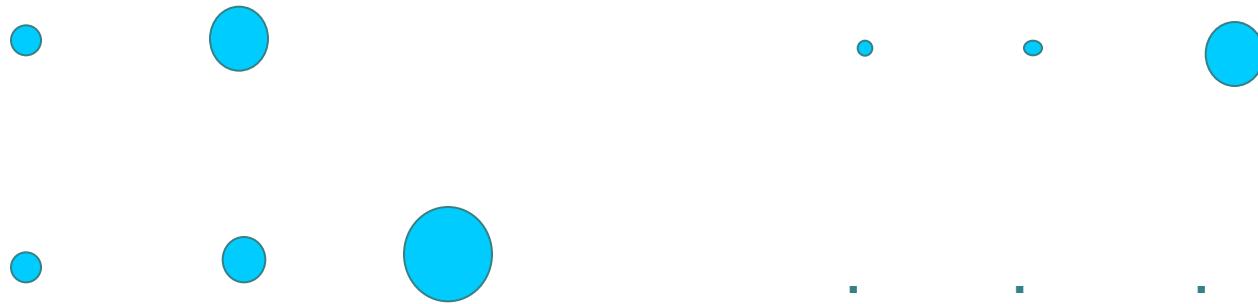
Antikvarkovi

Električni naboј		
Antidno	1/3	-2/3
Antistrani	1/3	-2/3
Antidolje	1/3	-2/3

Antivrh
Antišarmantni
Antigore

svaki antikvark:**R**, **B**, **G** 3 antiboje

MISTERIJ MASE



kvarkovi

leptoni

Razlog “bi moglo biti” postojanje
nove čestice, zovemo je “**Higgsov
boson**”

Veliki sudarač hadrona (LHC)



Veliki sudarač hadrona (LHC)

Large Hadron Collider

27 km circumference

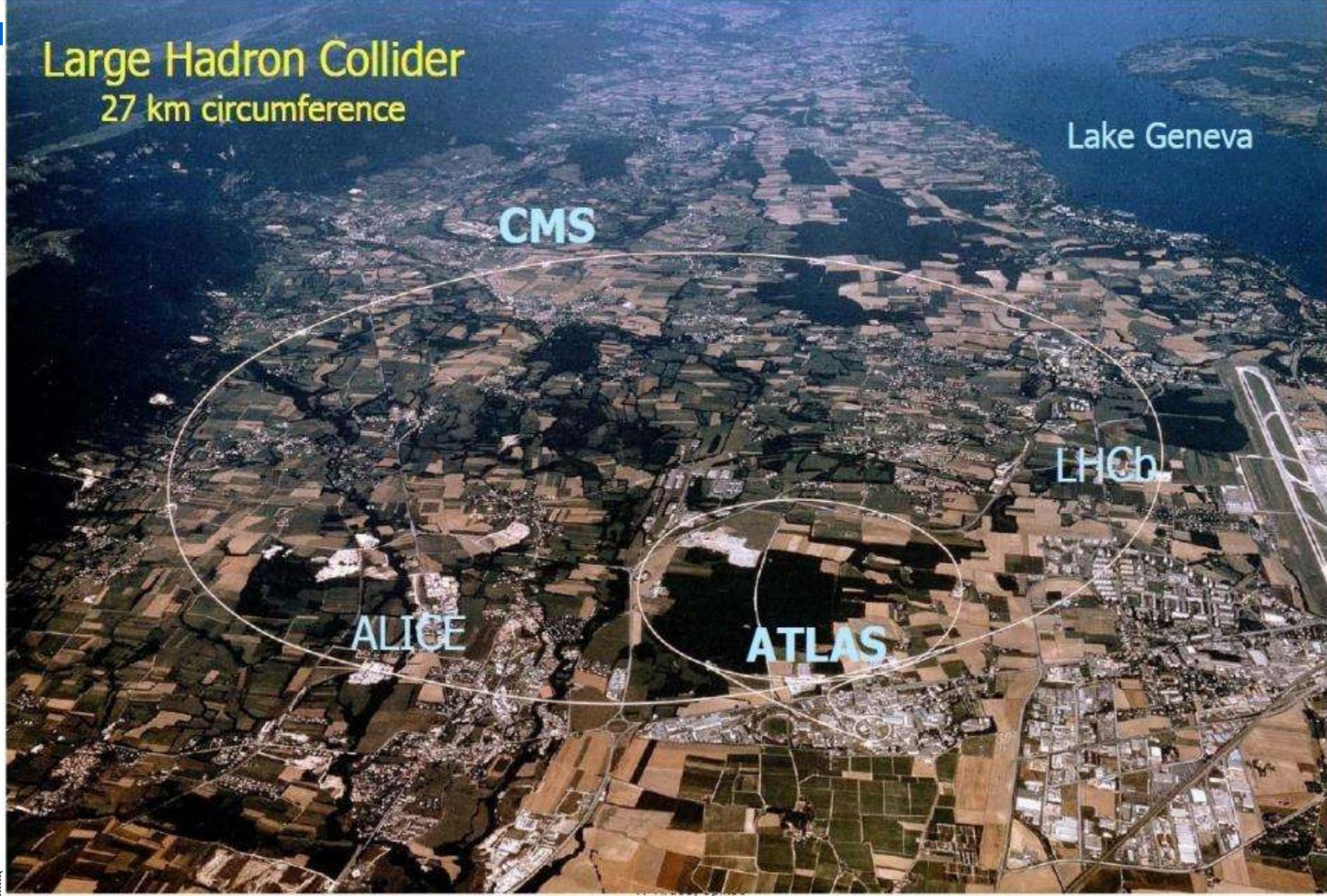
Lake Geneva

CMS

LHCb

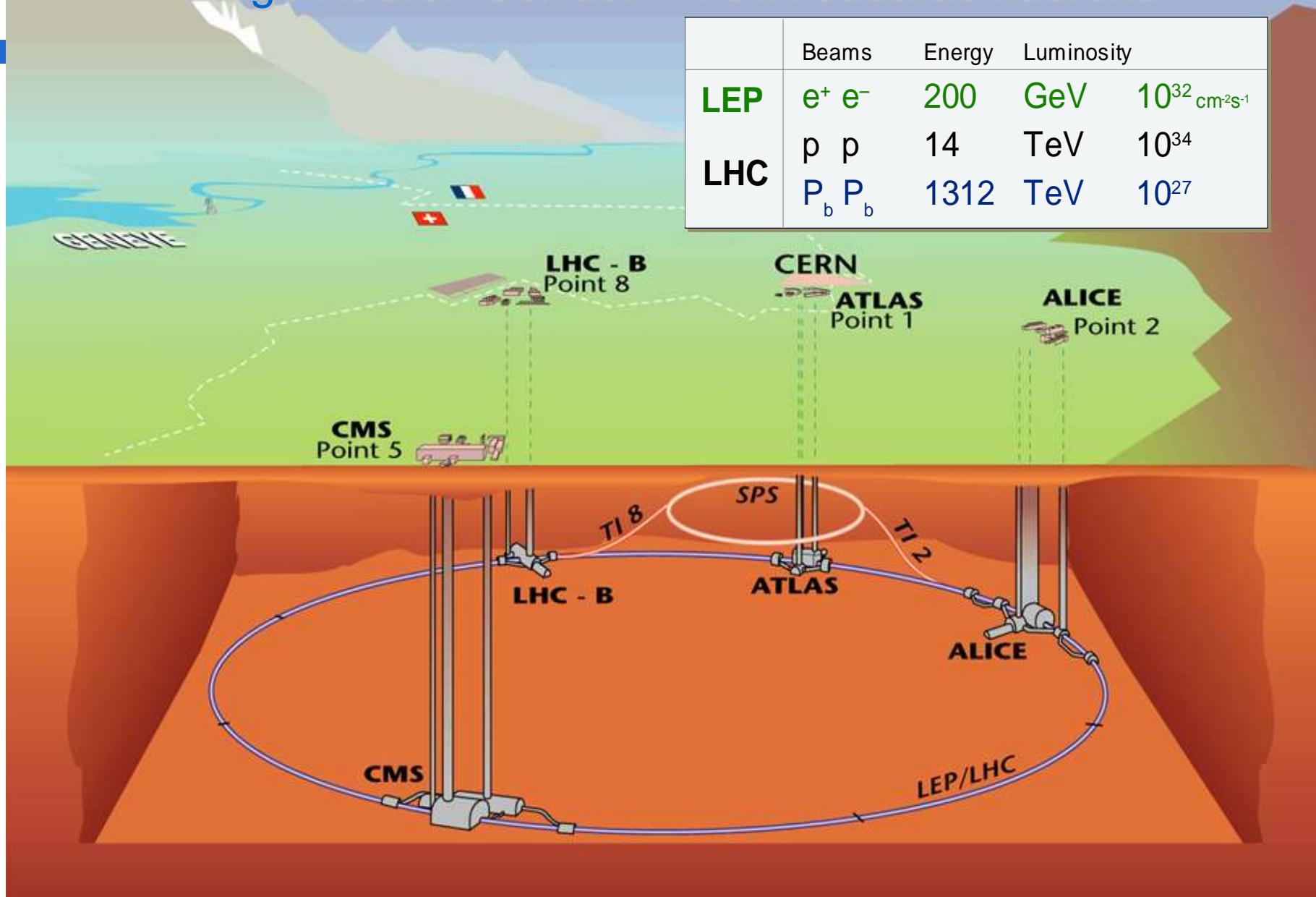
ALICE

ATLAS

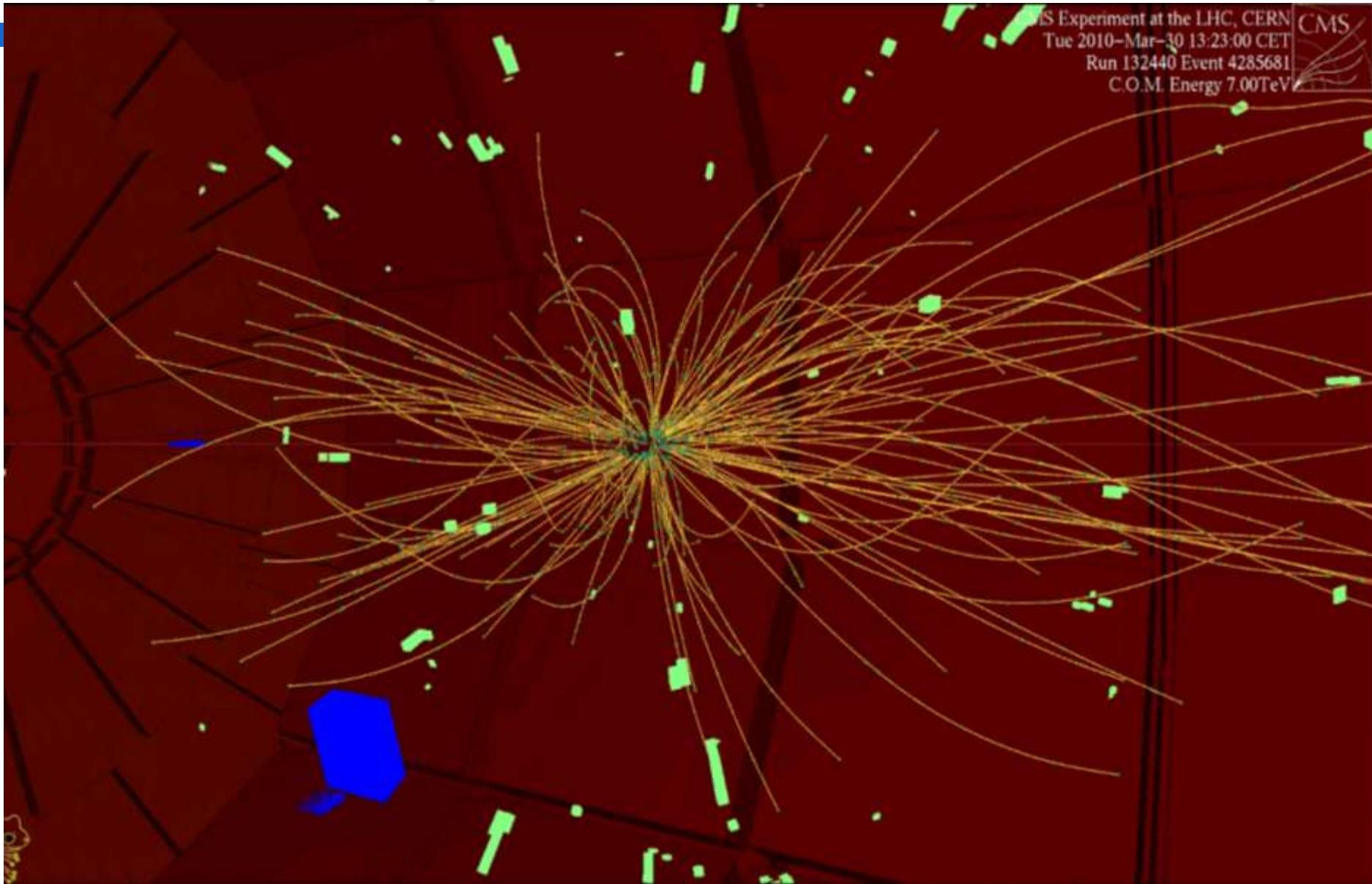


The Large Hadron Collider – Veliki sudarač hadrona

	Beams	Energy	Luminosity
LEP	$e^+ e^-$	200 GeV	$10^{32} \text{ cm}^{-2}\text{s}^{-1}$
LHC	$p p$	14 TeV	10^{34}
	$P_b P_b$	1312 TeV	10^{27}

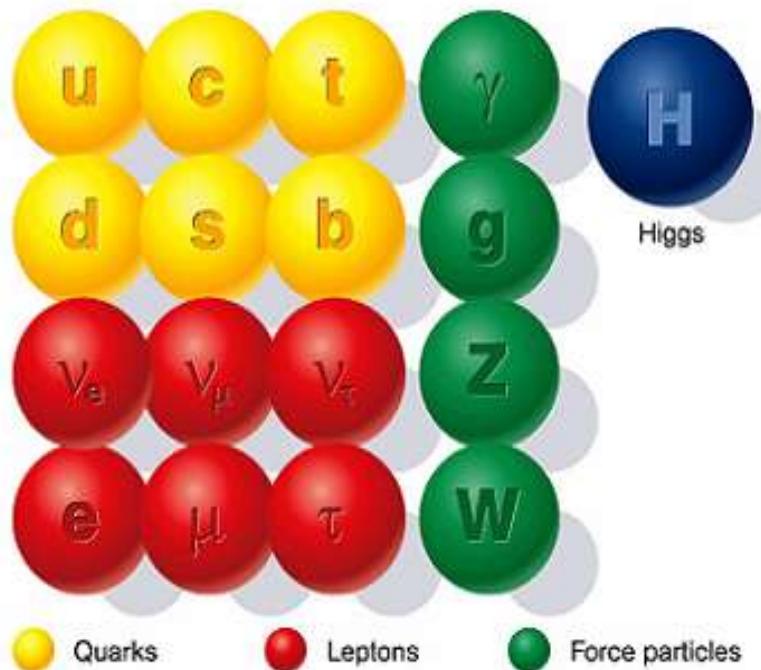


Prvi sudari protona u CMS detektoru



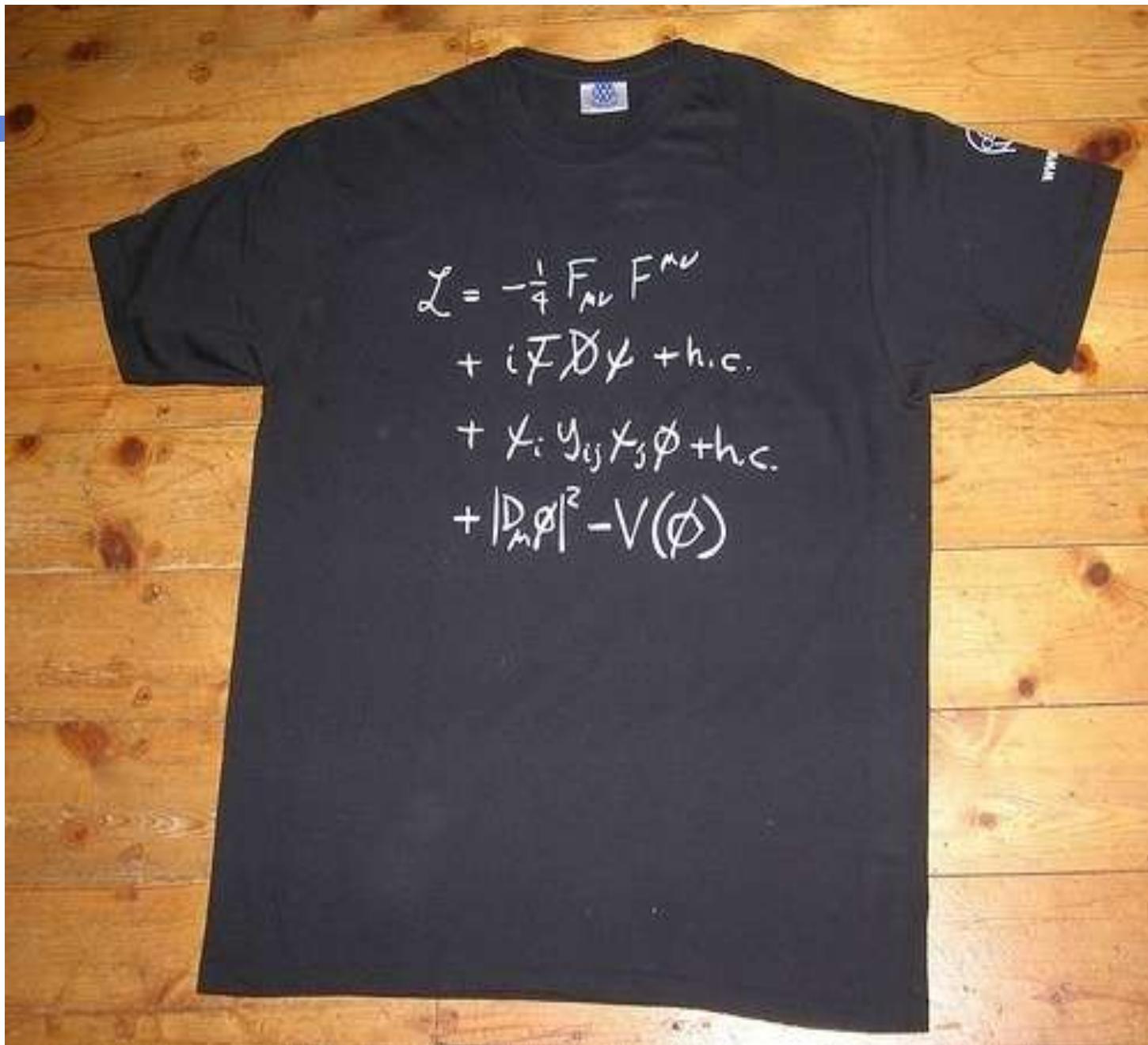
Standardni Model čestica i sila

Standard particles



$$\begin{aligned} \mathcal{L}_{GWS} = & \sum_f (\bar{\Psi}_f (i\gamma^\mu \partial_\mu - m_f) \Psi_f - e Q_f \bar{\Psi}_f \gamma^\mu \Psi_f A_\mu) + \\ & + \frac{g}{\sqrt{2}} \sum_i (\bar{a}_L^i \gamma^\mu b_L^i W_\mu^+ + \bar{b}_L^i \gamma^\mu a_L^i W_\mu^-) + \frac{g}{2c_w} \sum_f \bar{\Psi}_f \gamma^\mu (I_f^3 - 2s_w^2 Q_f - I_f^3 \gamma_5) \Psi_f Z_\mu + \\ & - \frac{1}{4} |\partial_\mu A_\nu - \partial_\nu A_\mu - ie(W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 - \frac{1}{2} |\partial_\mu W_\nu^+ - \partial_\nu W_\mu^+ + \\ & - ie(W_\mu^+ A_\nu - W_\nu^+ A_\mu) + ig' c_w (W_\mu^+ Z_\nu - W_\nu^+ Z_\mu)|^2 + \\ & - \frac{1}{4} |\partial_\mu Z_\nu - \partial_\nu Z_\mu + ig' c_w (W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 + \\ & - \frac{1}{2} M_\eta^2 \eta^2 - \frac{g M_\eta^2}{8 M_W} \eta^3 - \frac{g'^2 M_\eta^2}{32 M_W} \eta^4 + |M_W W_\mu^+ + \frac{g}{2} \eta W_\mu^+|^2 + \\ & + \frac{1}{2} |\partial_\mu \eta + i M_Z Z_\mu + \frac{ig}{2c_w} \eta Z_\mu|^2 - \sum_f \frac{g}{2} \frac{m_f}{M_W} \bar{\Psi}_f \Psi_f \eta \end{aligned}$$

$$\begin{aligned}\mathcal{L} = & -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \\& + i \bar{\psi} D^\mu \psi + h.c. \\& + Y_i Y_{ij} Y_j \phi + h.c. \\& + |D_\mu \phi|^2 - V(\phi)\end{aligned}$$



Higgsov mehanizam

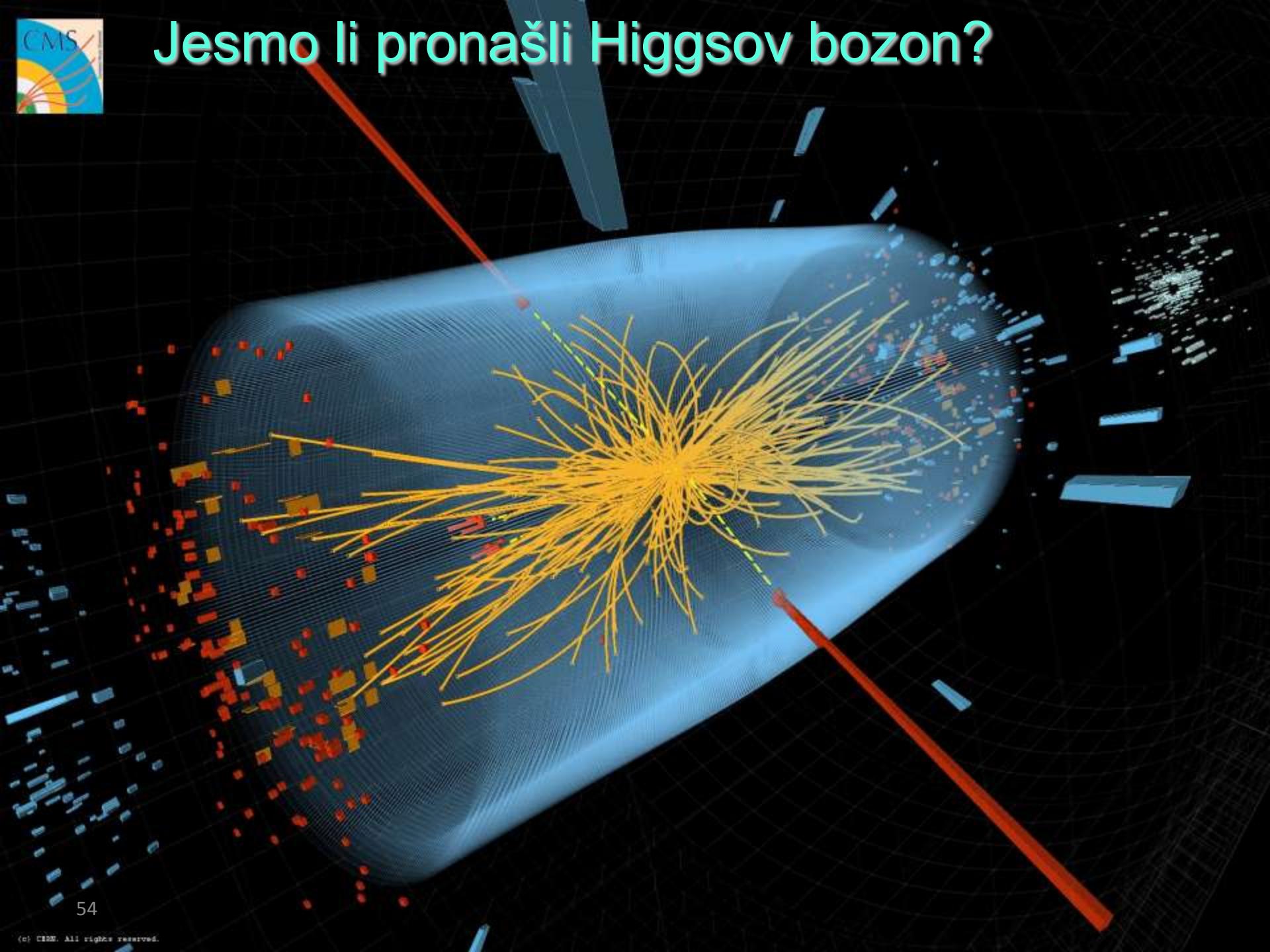


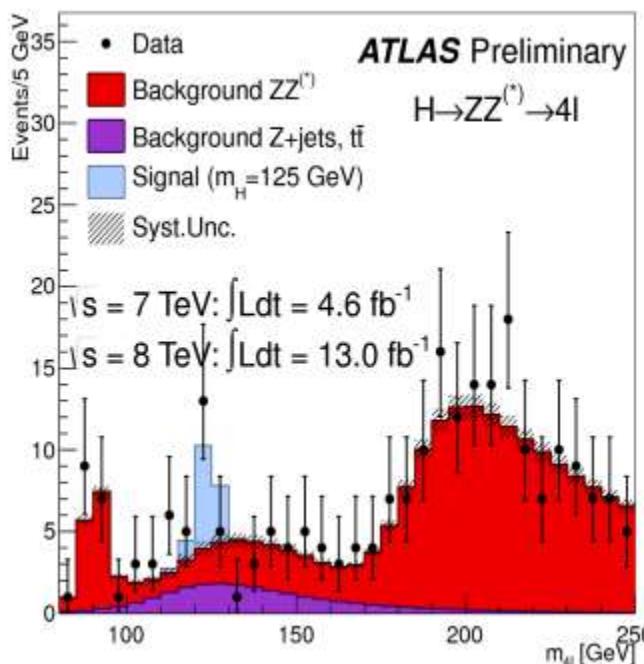
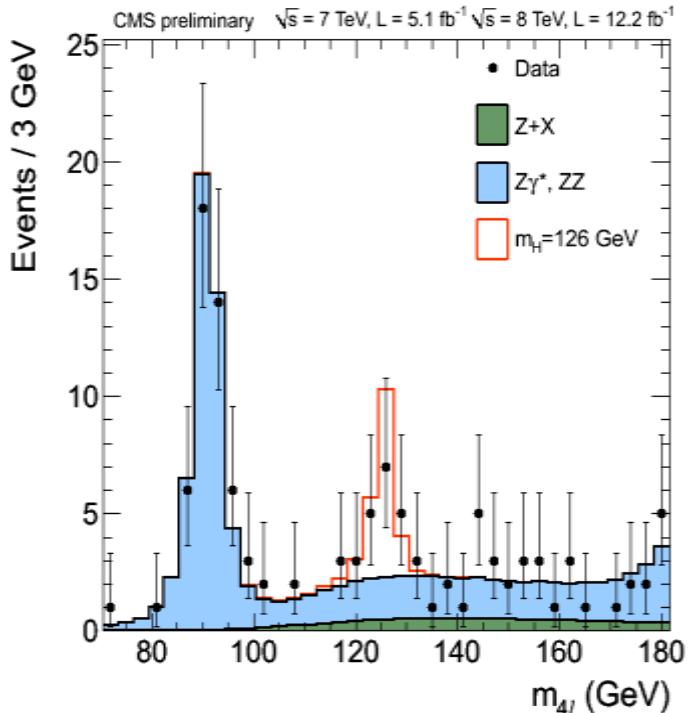
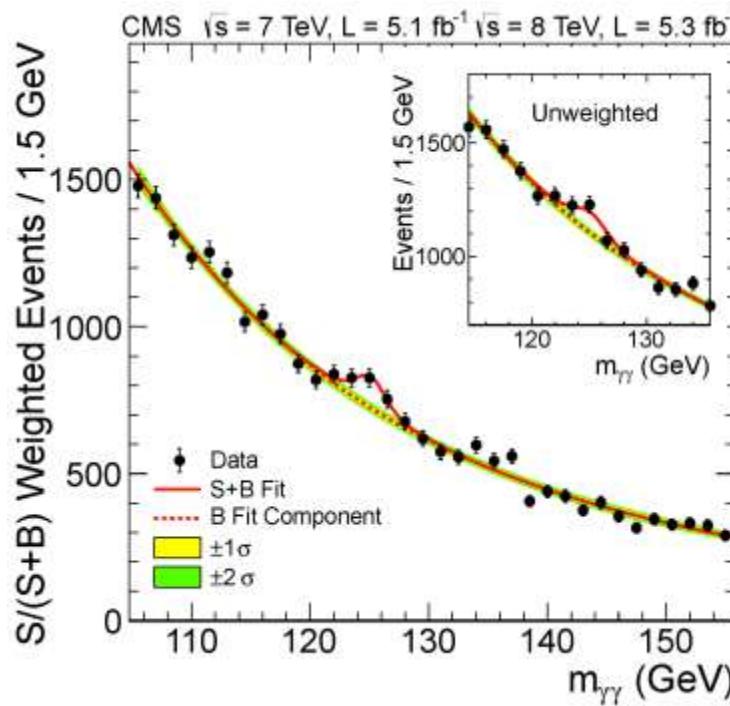
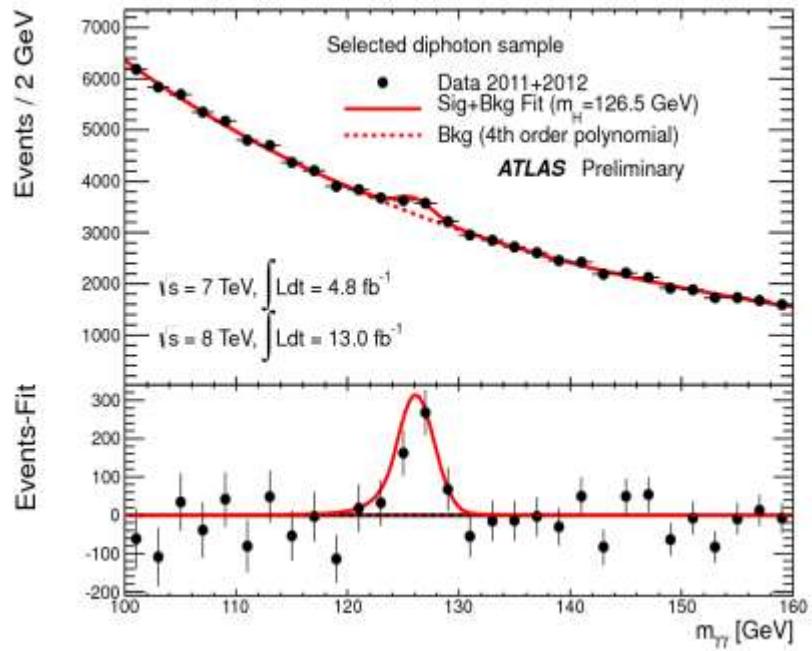
Higgsov bozon





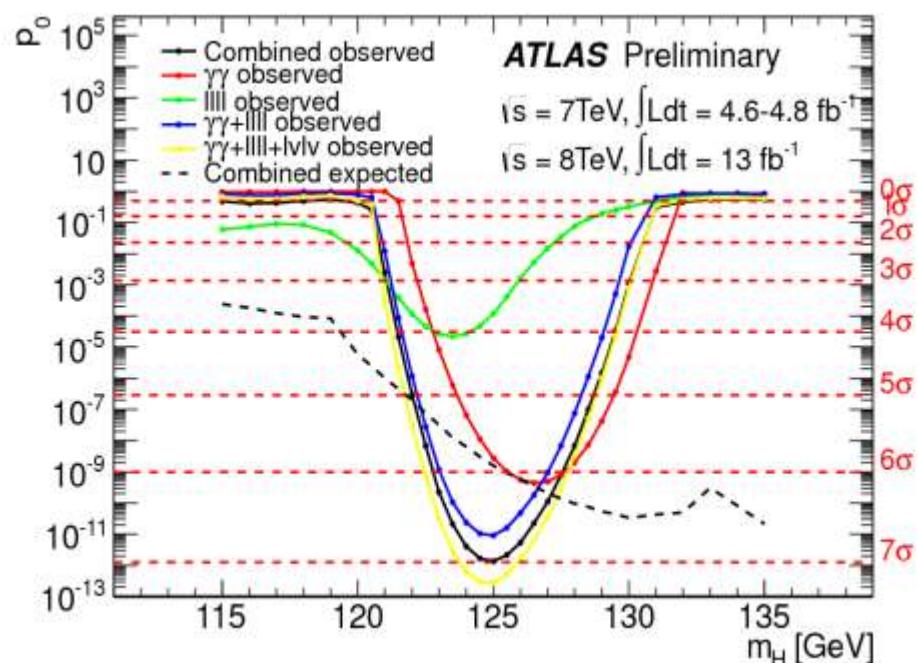
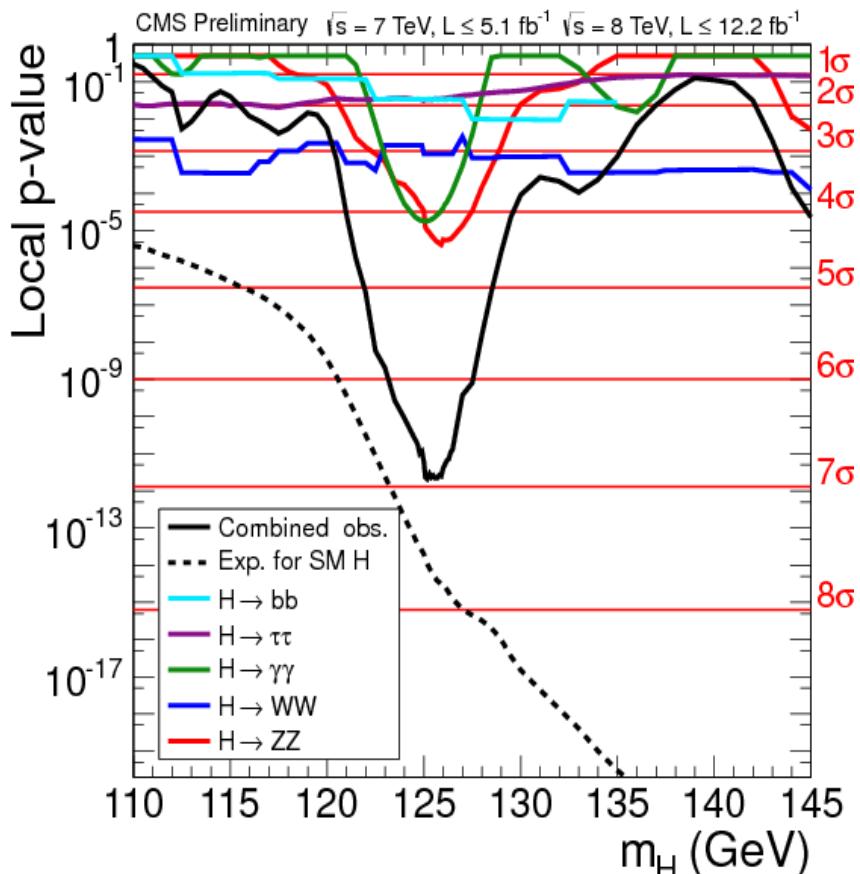
Jesmo li pronašli Higgsov bozon?



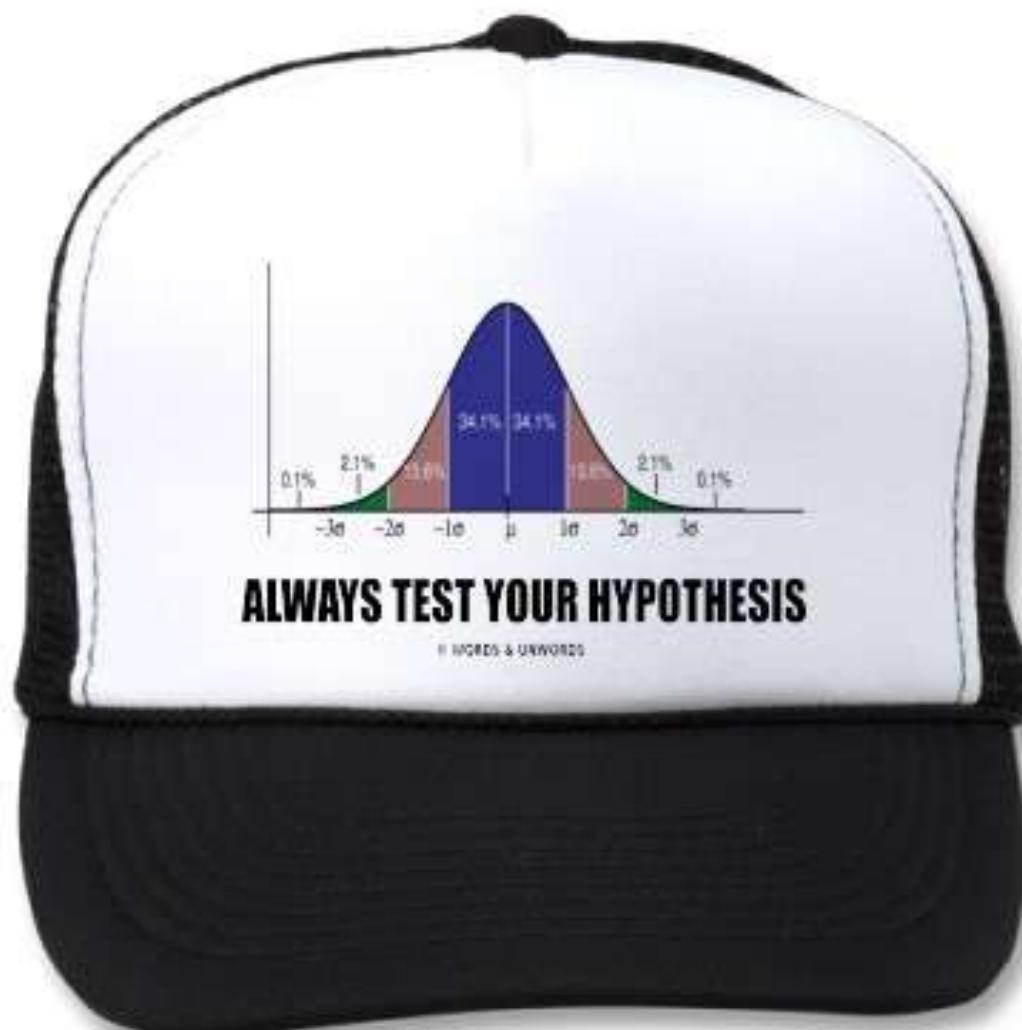


Višak podataka u svima kanalima

Na istom mjestu



Vrlo vjerojatno je to Higgsov bozon ...



Je li vrijeme iluzija?

